

1 A GeV ^{238}U on proton. Transition from fission to light evaporation residues. Data and calculation.

Paper:

**"Light Nuclides Produced in the Proton-Induced
Spallation of ^{238}U at 1 GeV"**

accepted in Phys. Rev. C

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PART I:

DATA ANALYSIS

THE EXPERIMENTAL SET-UP for LIGHT RESIDUES

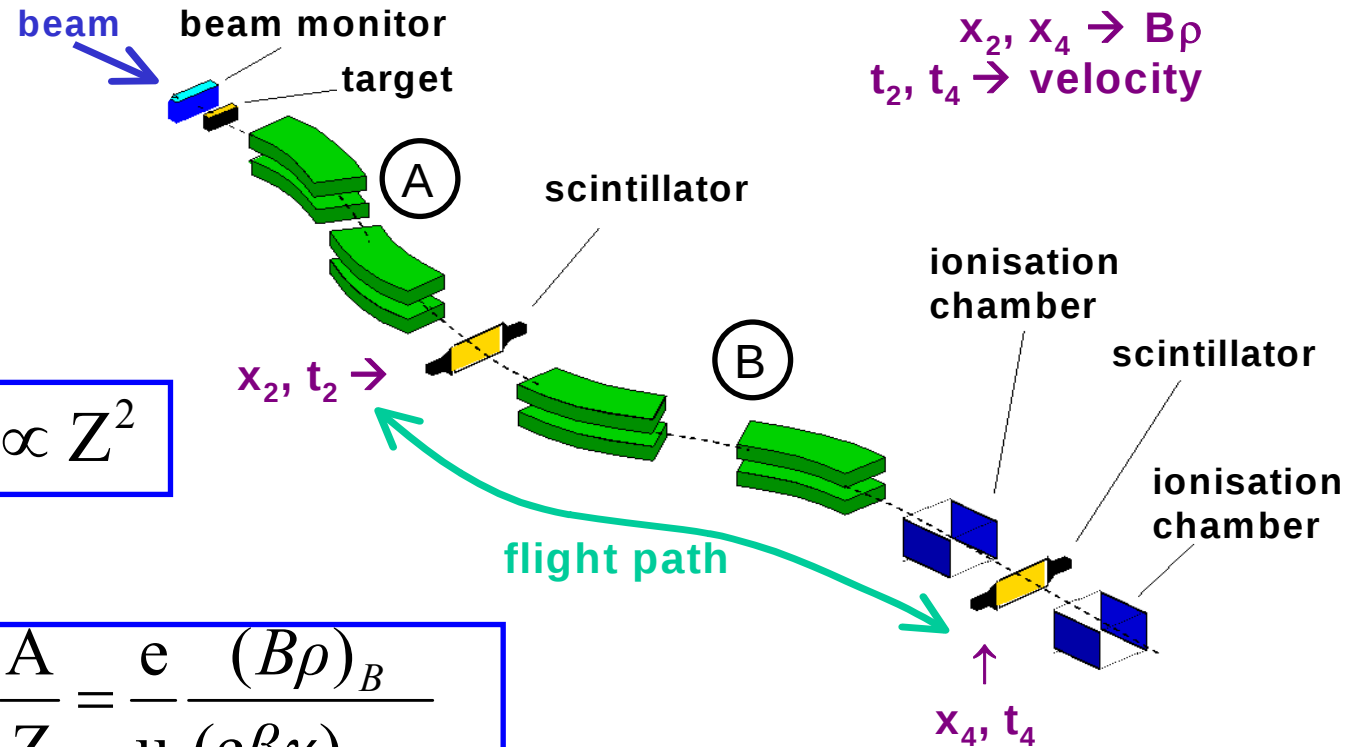
Analysis in
the range:
 $7 \leq Z \leq 37$

Z from IC: $\Delta E \propto Z^2$

A/Z from time
and position: $\frac{A}{Z} = \frac{e}{u} \frac{(B\rho)_B}{(c\beta\gamma)_{TOF}}$

velocity from magnetic
rigidity:

$$\beta\gamma c = \frac{e}{u + \delta u} \cdot \frac{Z}{A} \cdot (B\rho)_A$$



CHARGE RESOLUTION

$$\Delta E_{theory} \cong \frac{Z^2}{f(\nu)}$$

The MUSIC signals depended on two other additional factors:

- on the distance of the ions from the anode, i.e. the x_4 position, due to a recombination effect, leading to a loss of free electrons in the gas
- on the density of the gas in the IC, which was observed to vary with the time at which ΔE was measured

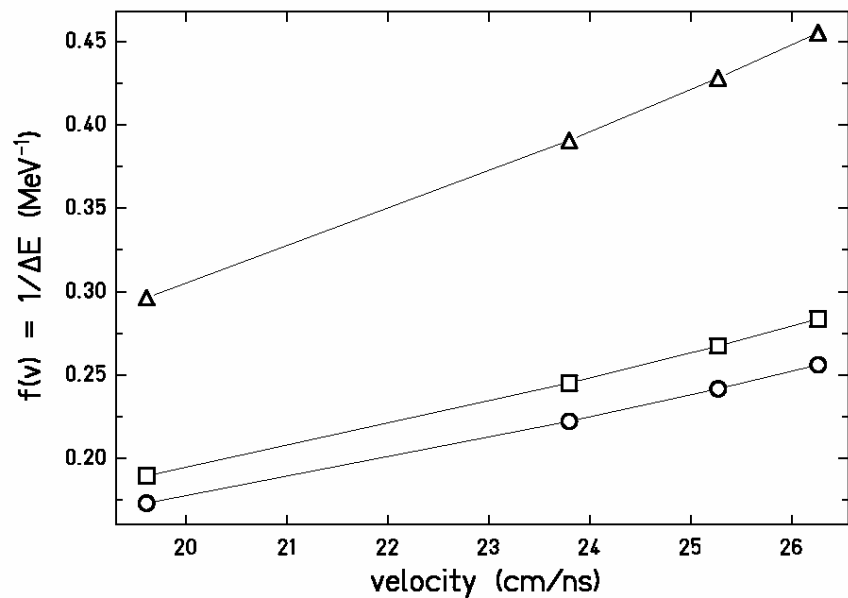
$$\Delta E_{measured} \cong \frac{Z^2}{f(\nu)} \cdot p(t) \cdot g(x_4)$$

$$Z^2 \propto \Delta E_{measured} \cdot \frac{f(\nu)}{p(t) \cdot g(x_4)}$$

(we assumed that the variables ν , t , x_4 , are separable)

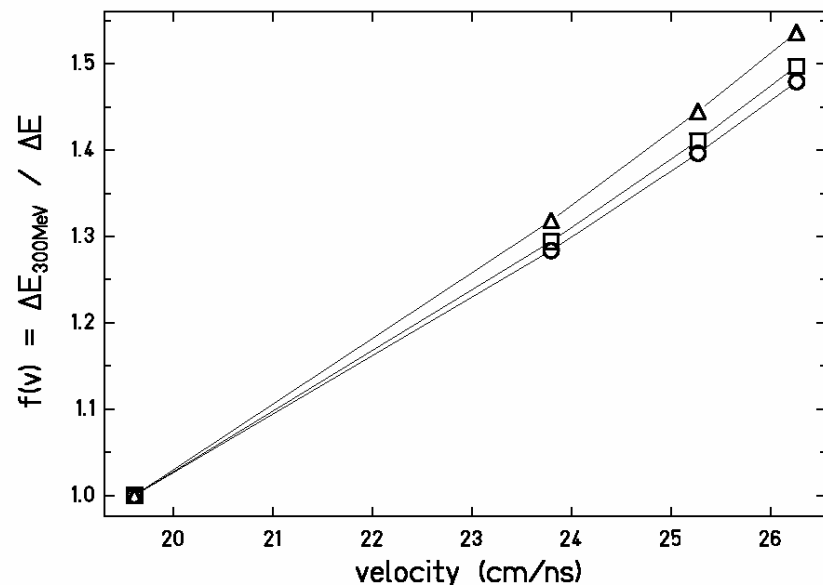
CHARGE RESOLUTION

Correction for velocity dependence



○ U
□ Pb
△ Sn

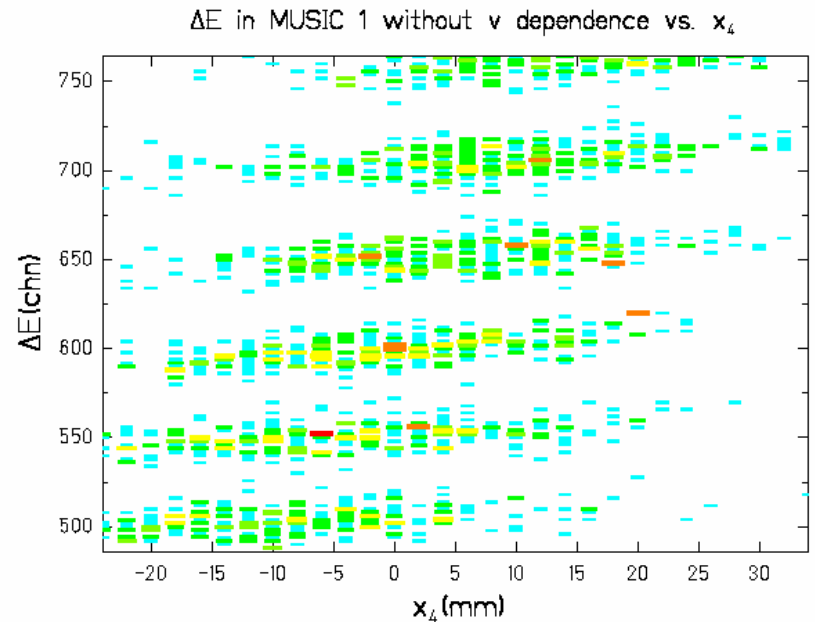
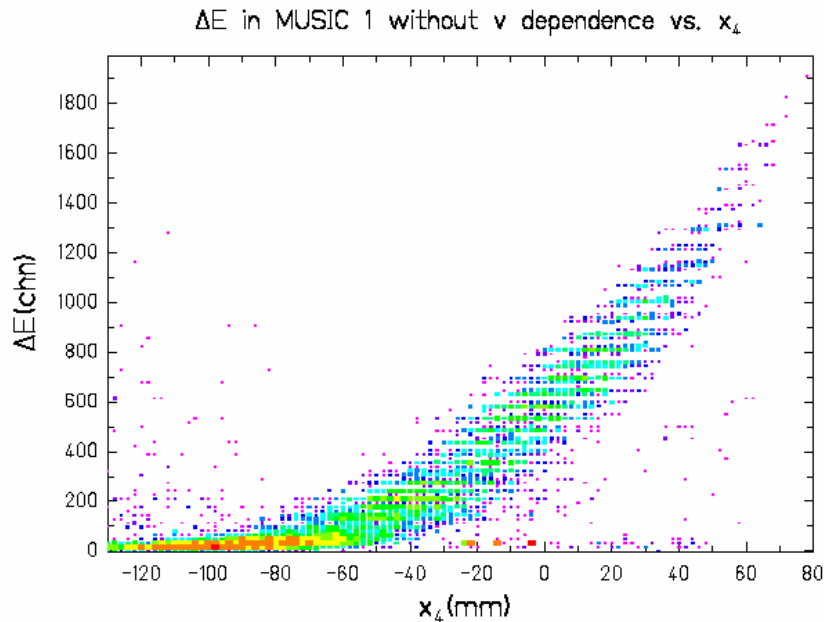
$$f(v) \propto \Delta E_{theory}^{300MeV} / \Delta E_{theory}$$



○ U
□ Pb
△ Sn

CHARGE RESOLUTION

Correction for position dependence



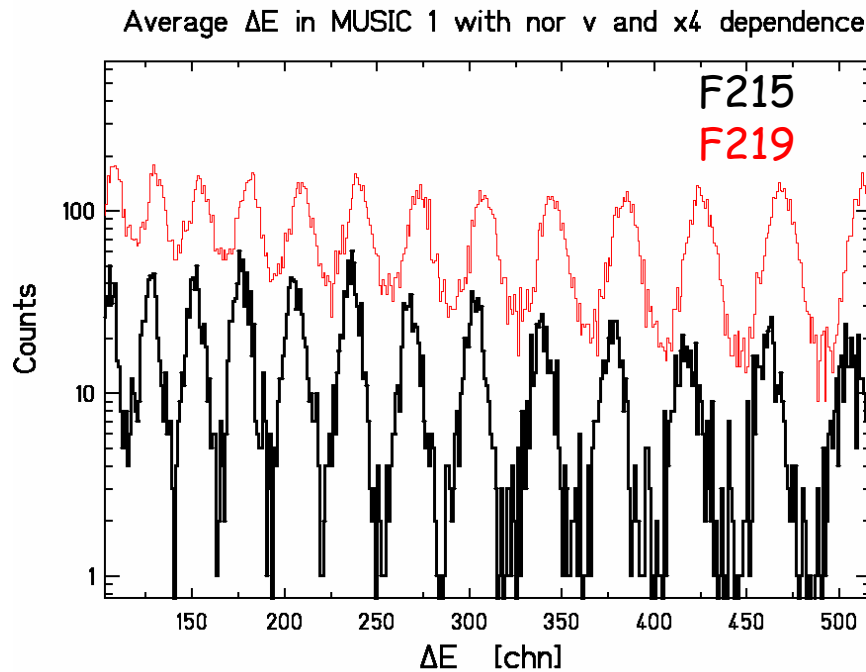
$$\Delta E_{\text{corr}} = \frac{\Delta E_{\text{meas}}}{g(x_4)} = \frac{\Delta E_{\text{meas}}}{e^{\lambda \cdot x_4}} = \Delta E_{\text{meas}} \cdot e^{-\lambda \cdot x_4}$$

$$\lambda = 5 \cdot 10^{-4}$$

The exponential dependence is due to the fact that the recombination of the electrons follows an *absorption law*, i.e. the number of electrons traversing a certain distance dx₄ is reduced each time by the same percentage value.

CHARGE RESOLUTION

Correction for pressure dependence



$$p(t)_{\Delta E1} = -4.44 + 0.0479 \cdot (\text{filenumber}) - 1.05 \cdot 10^{-4} \cdot (\text{filenumber})^2$$

MASS RESOLUTION

$$\frac{A}{Z} = \frac{e}{u} \frac{(B\rho)_B}{\beta_{TOF} \gamma_{TOF} c}$$

The A/Z depends on two measured quantities:

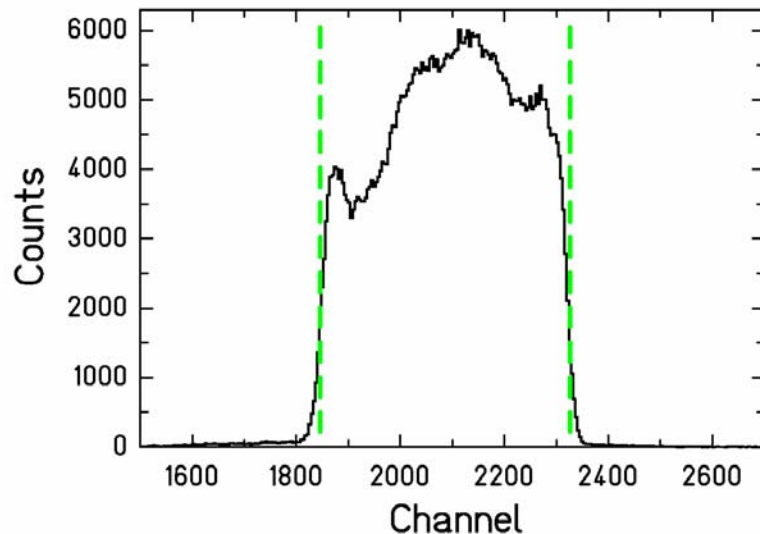
- the magnetic rigidity in the second stage $(B\rho)_B$
- the velocity from TOF

MASS RESOLUTION

The magnetic rigidity in the second stage $(B\rho)_B$

$$\Delta(B\rho)_{AB} = (B\rho)_A - (B\rho)_B \quad (B\rho)_B = (B\rho)_{A0} \cdot \left(1 + \frac{x_2}{D_2} \right) - \Delta(B\rho)_{AB}$$

$\Delta(B\rho)_{AB}/(B\rho)_B$
from 0.2% to 1%



$$x_2(mm) = b_2 - a_2 \cdot SCI_2(channel)$$

MASS RESOLUTION

The velocity from ToF measurement

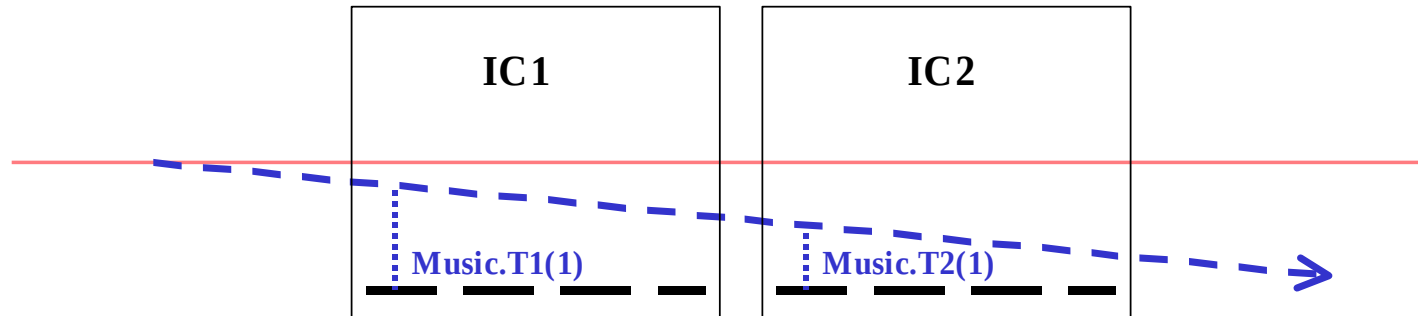
$$v_{TOF} = \frac{s}{ToF} = \frac{s_0 \cdot (1 + c_\alpha \alpha_x) + \Delta s}{T_0 - ToF^*} \quad \text{with} \quad \begin{cases} \Delta s = d_1 x_2 + d_2 x_2^2 \\ ToF^* = \frac{ToF_L^* \cdot \alpha_L + ToF_R^* \cdot \alpha_R}{2} \end{cases}$$

The flight path, s , differs from s_0 for two reasons:

- $(1 + c_\alpha \alpha_x)$ = fragments acquire a transversal momentum $\rightarrow$ enter the 2nd section of FRS with different angles
- ΔS = fragments with different $B\rho$ enter the 2nd section of FRS with different x_2 -positions $\rightarrow$ different paths along FRS

MASS RESOLUTION

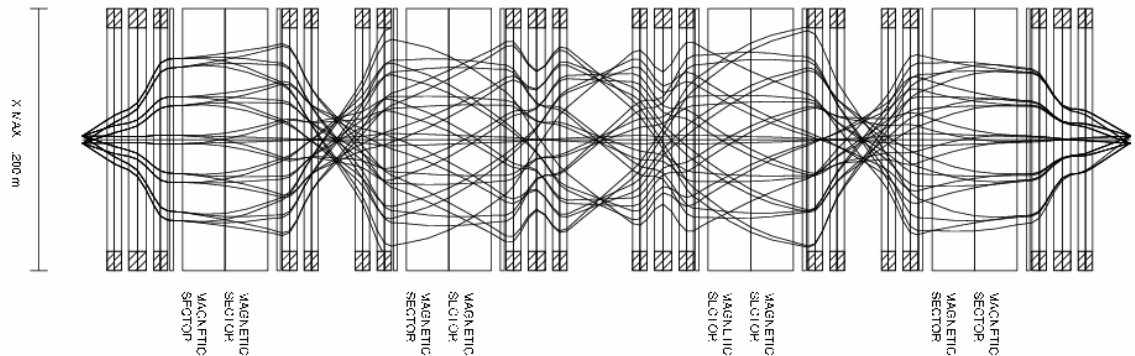
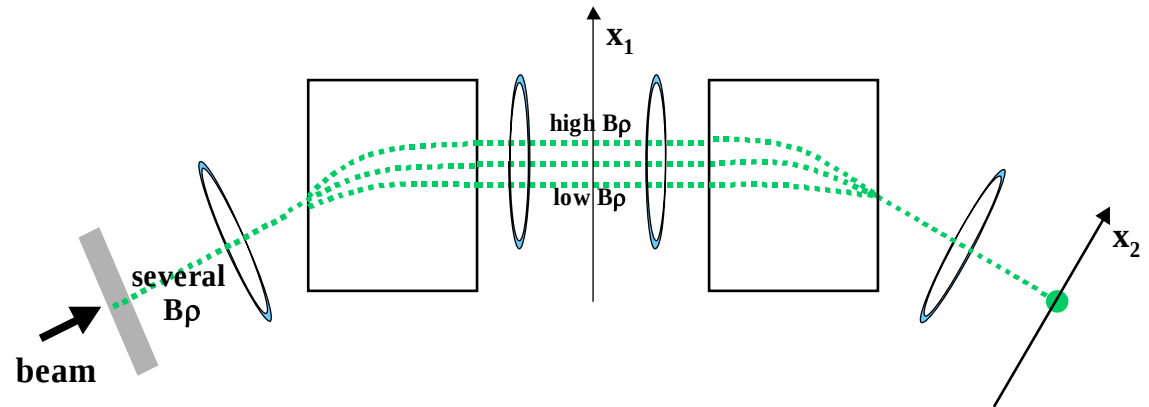
$(1+c_\alpha \alpha_x)$ = fragments acquire a transversal momentum $\rightarrow$ enter the 2nd section of FRS with different angles



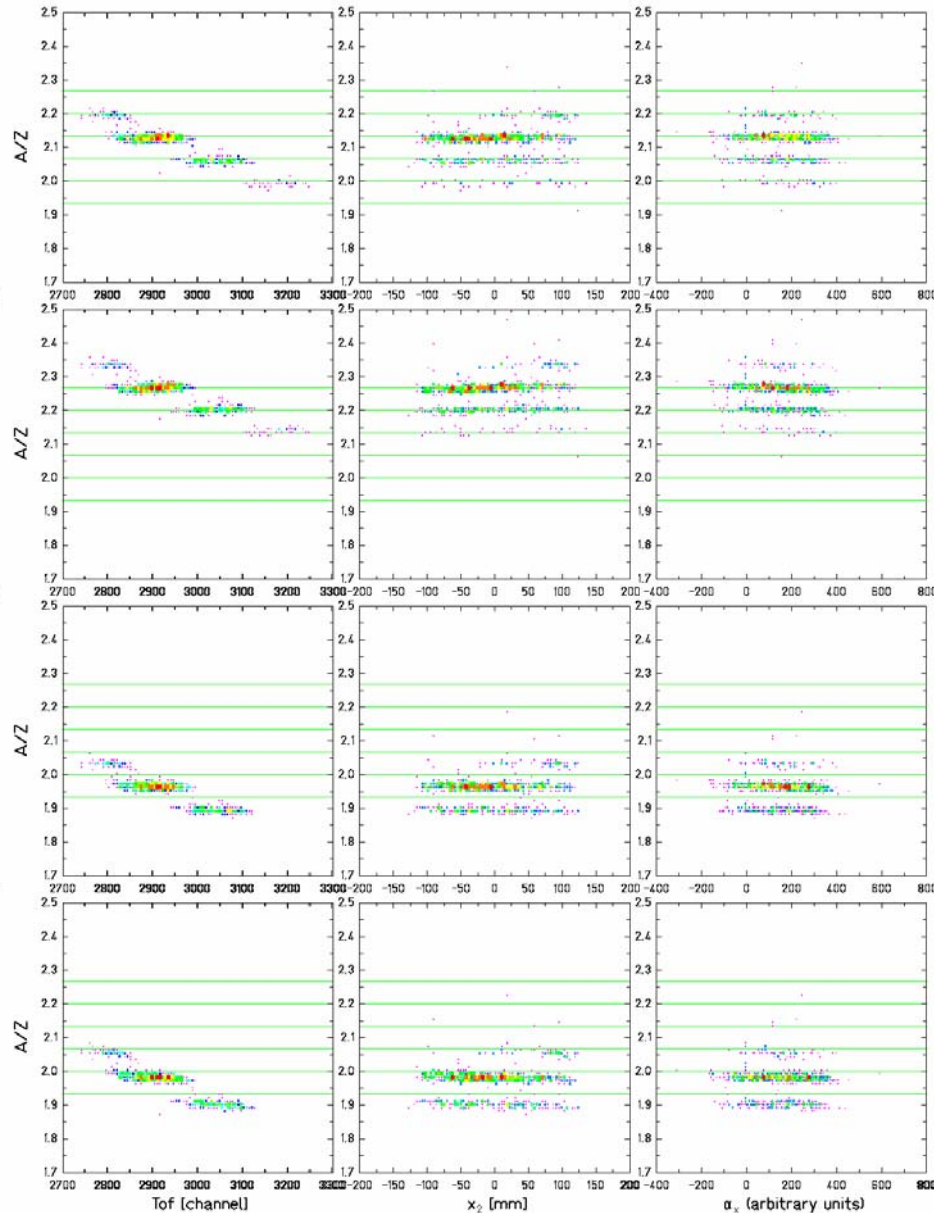
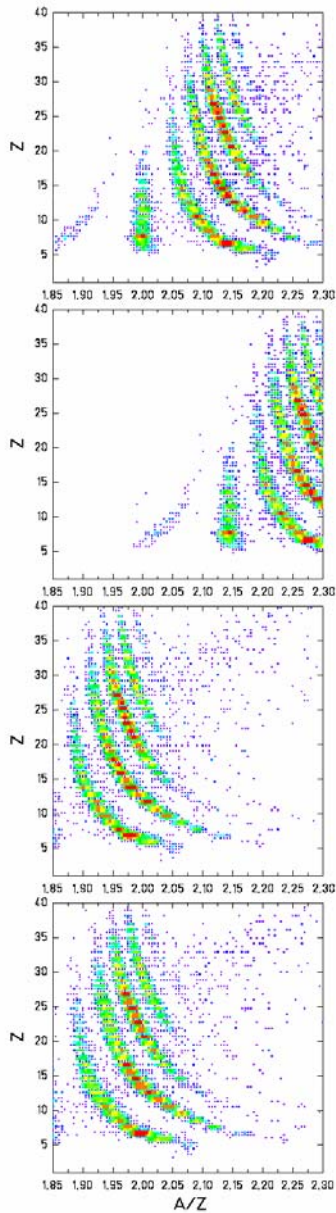
MASS RESOLUTION

ΔS = fragments with different $B\rho$ enter the 2nd section of FRS with different x_2 -positions $\rightarrow$ different paths along FRS

Linear + quadratic term



A check for determining the good parameters

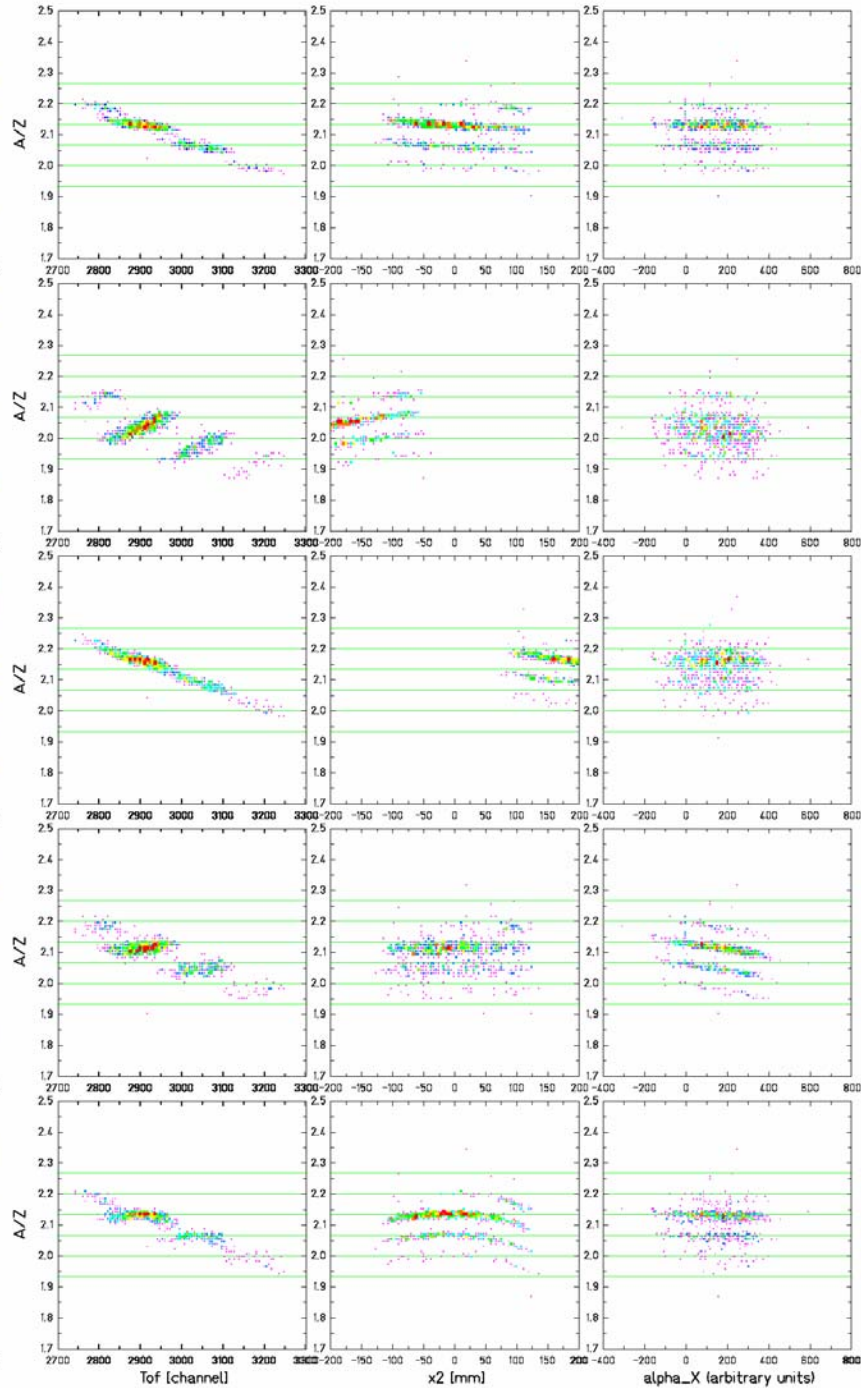
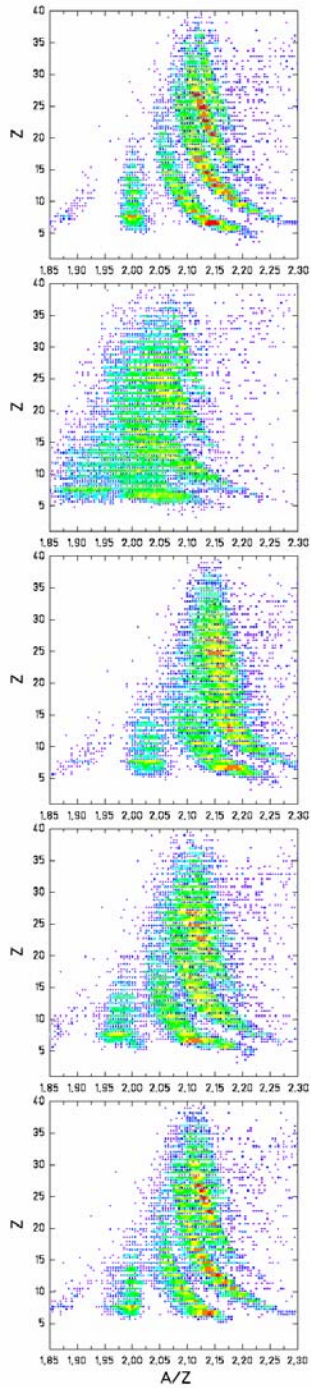


OK

wrong T_0

wrong s_0

wrong α_R



wrong D_2

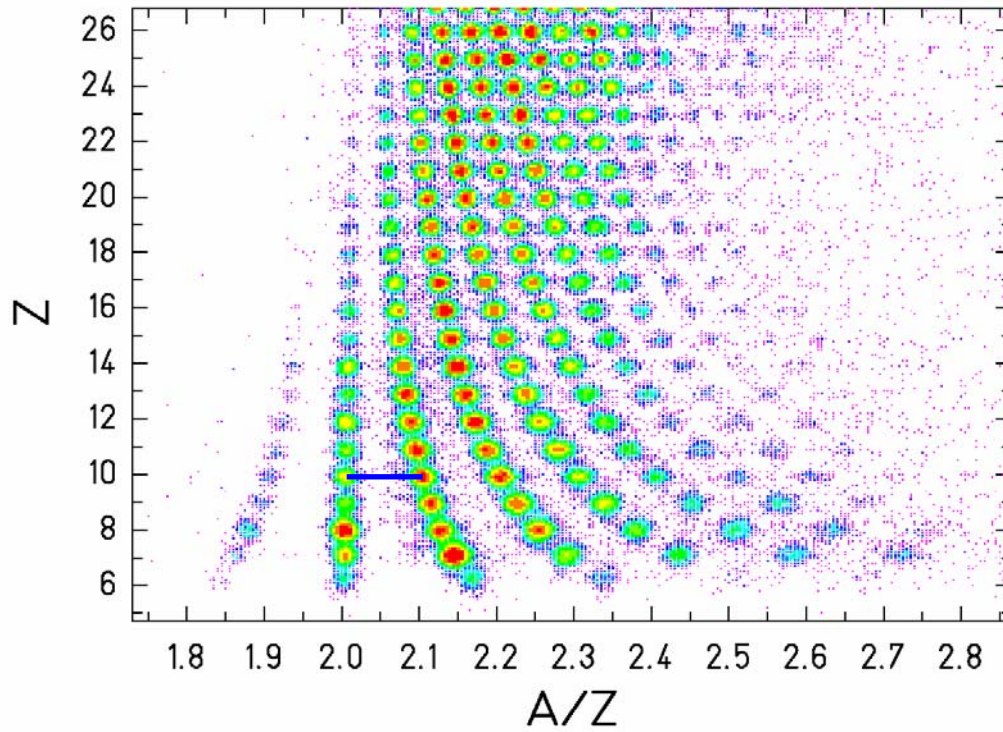
wrong a_2

wrong b_2

wrong c_α

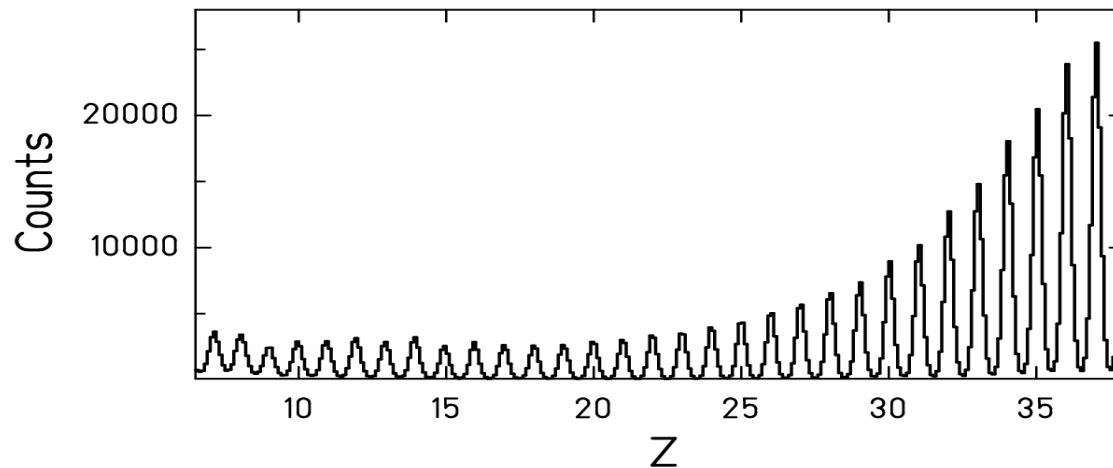
wrong d_2

RESOLUTION AND IDENTIFICATION

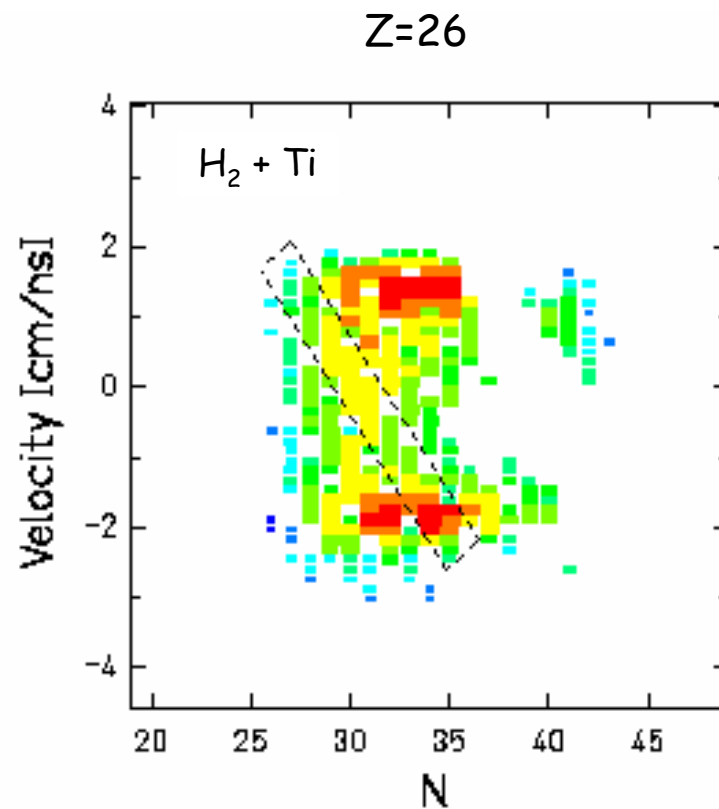
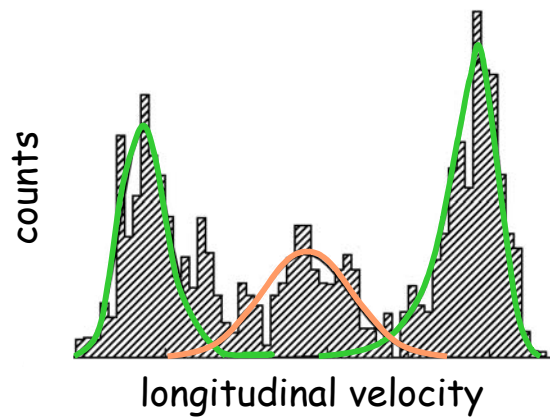
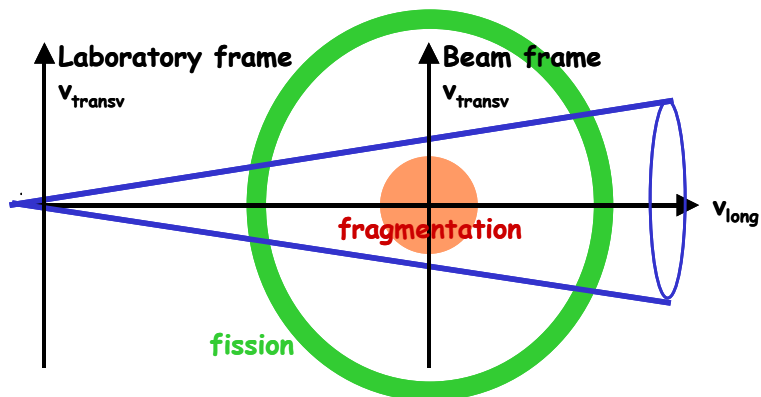


$$Z = \frac{-a_1 \pm \sqrt{a_1^2 - 4a_2(a_0 - \Delta E)}}{2a_2} + 4$$

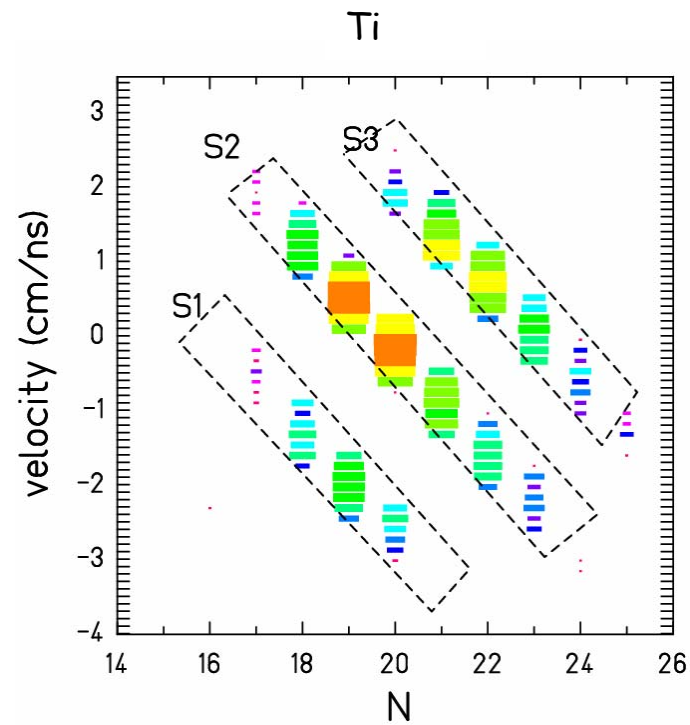
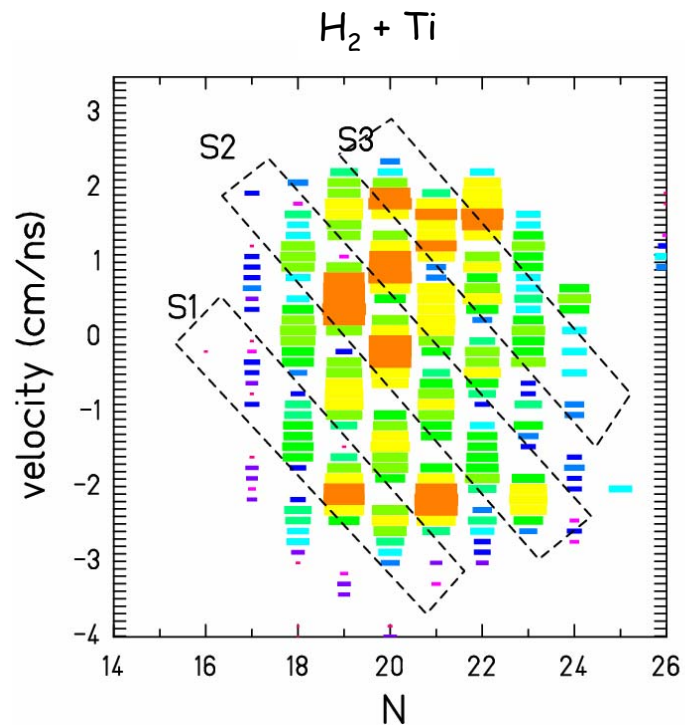
$$P = \frac{1}{((A+1)/Z) - (A/Z)} = Z$$



DETERMINATION OF THE VELOCITY



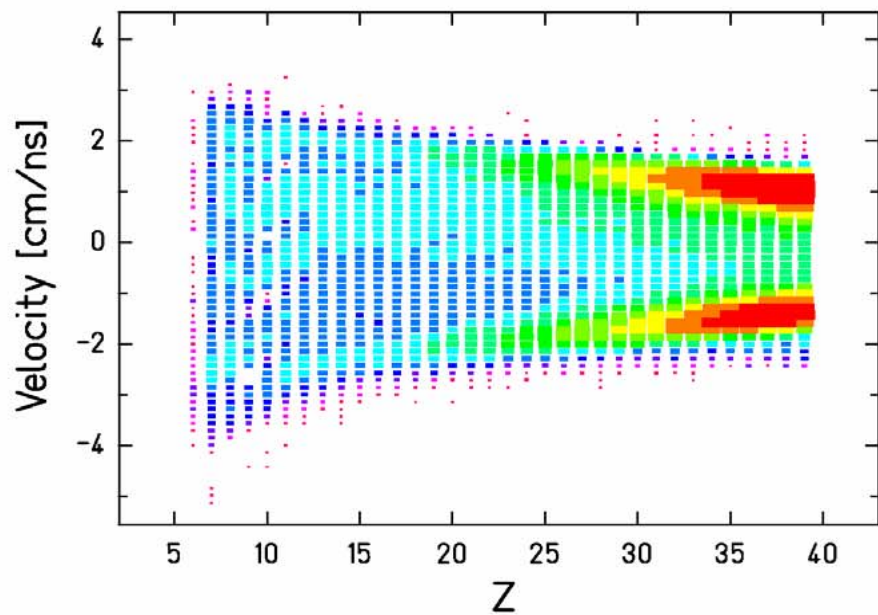
DETERMINATION OF THE VELOCITY



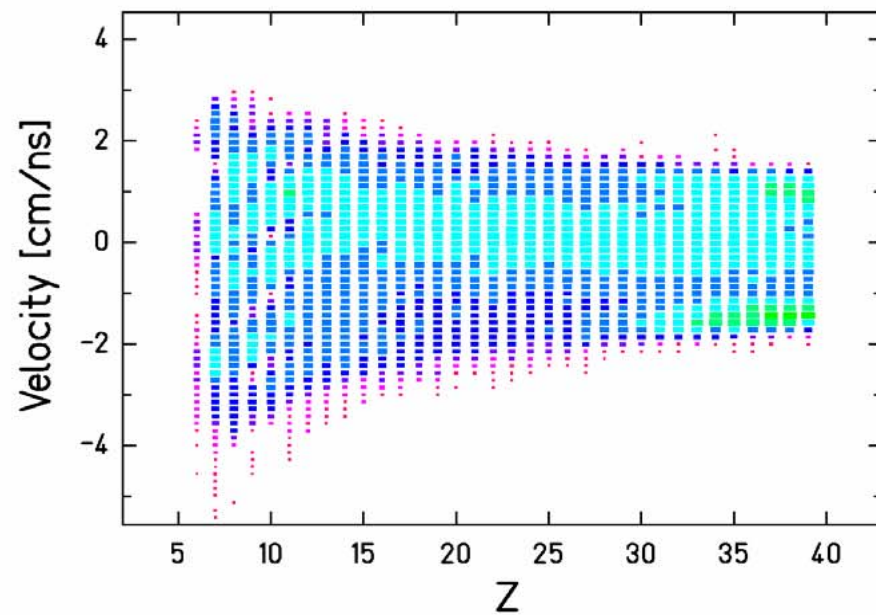
Too few settings for the dummy target!

VELOCITY SPECTRA FOR ALL RESIDUES

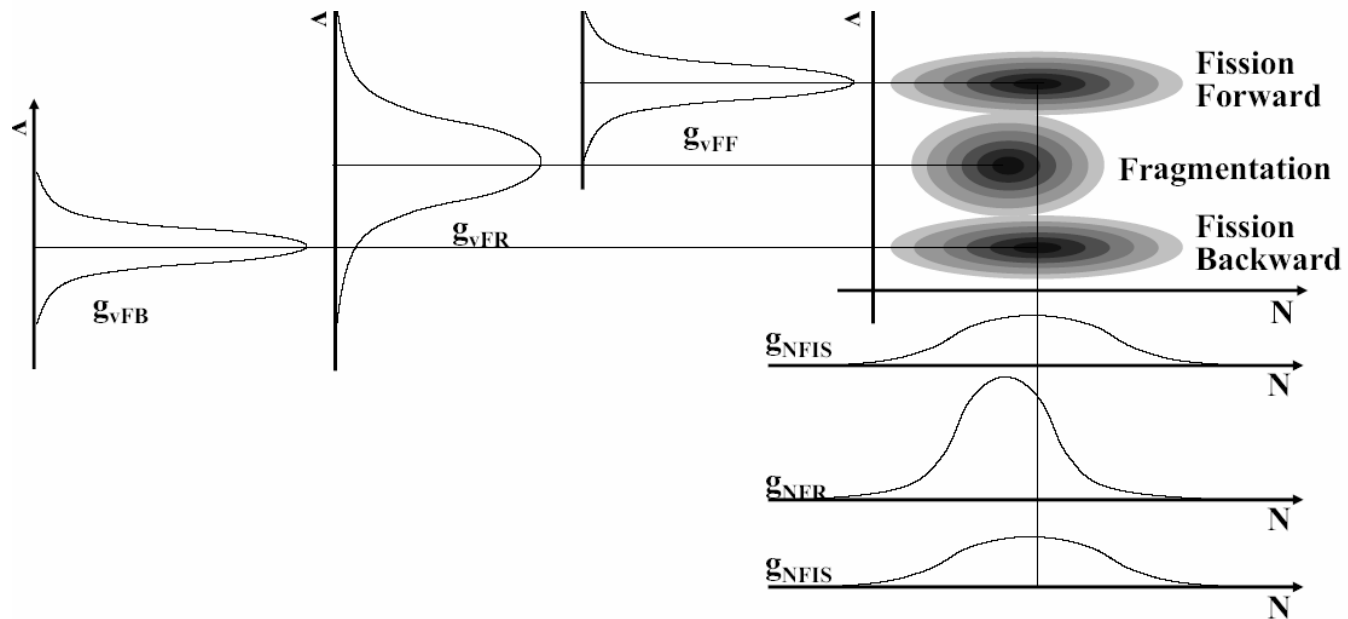
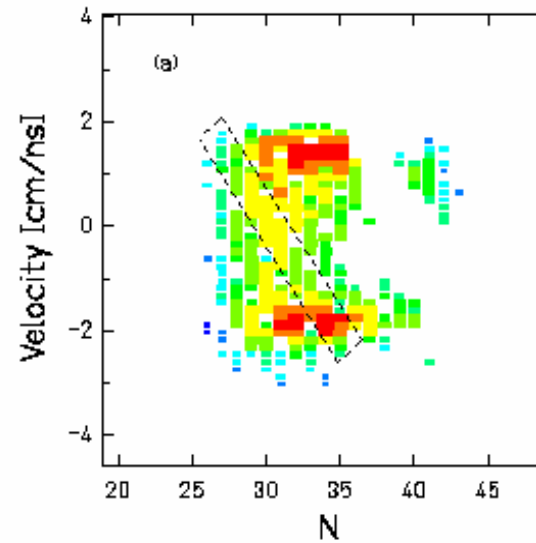
hydrogen+titanium



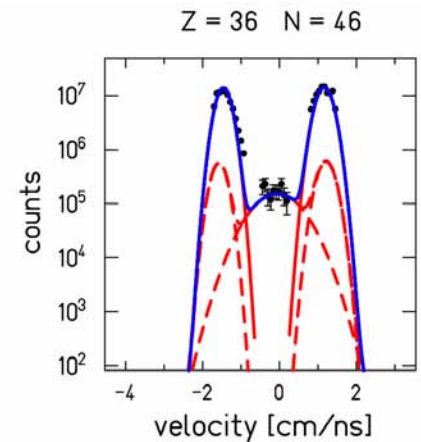
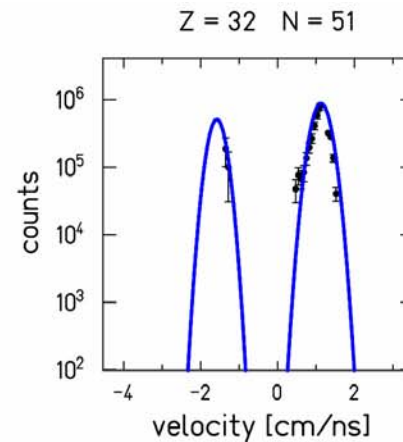
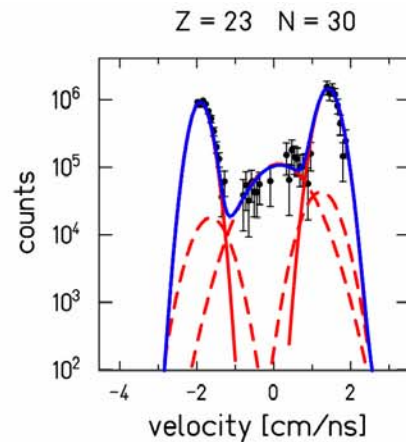
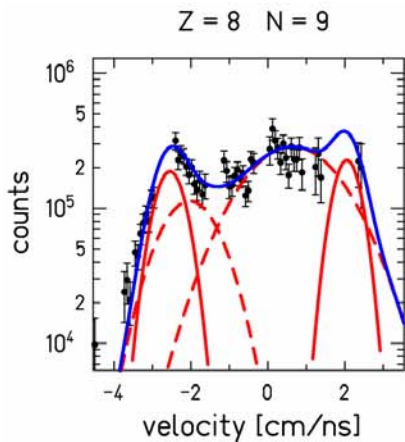
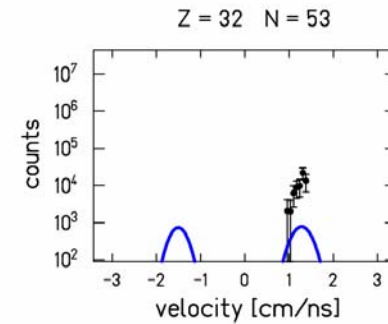
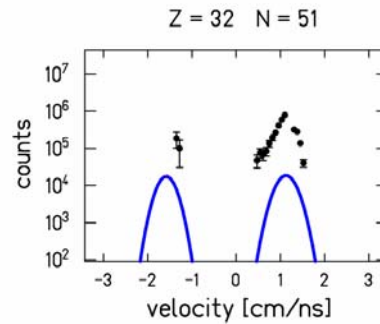
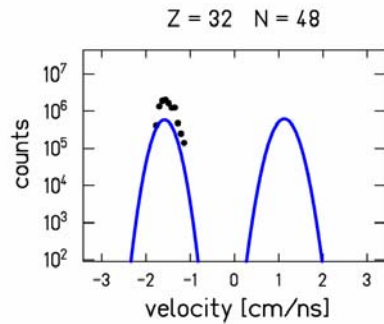
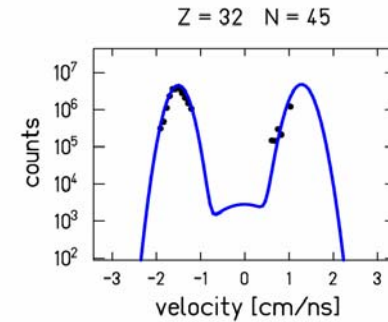
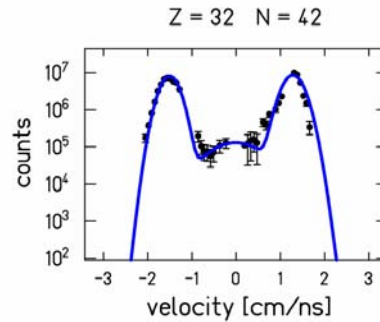
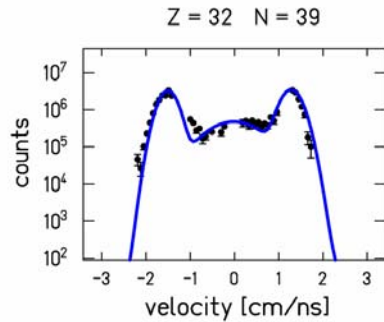
titanium alone



THE GLOBAL FIT



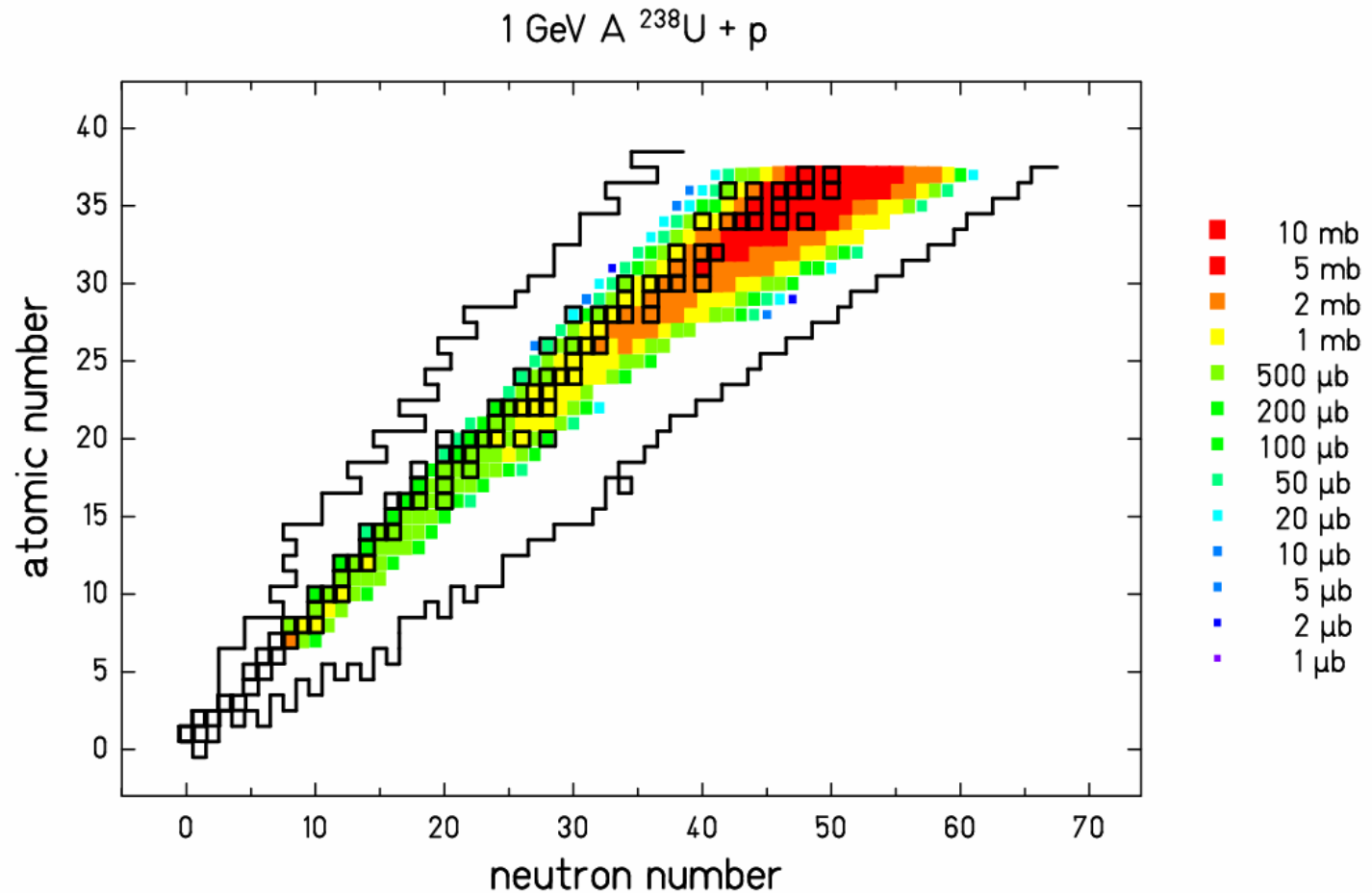
THE GLOBAL FIT AND THE INDIVIDUAL FIT



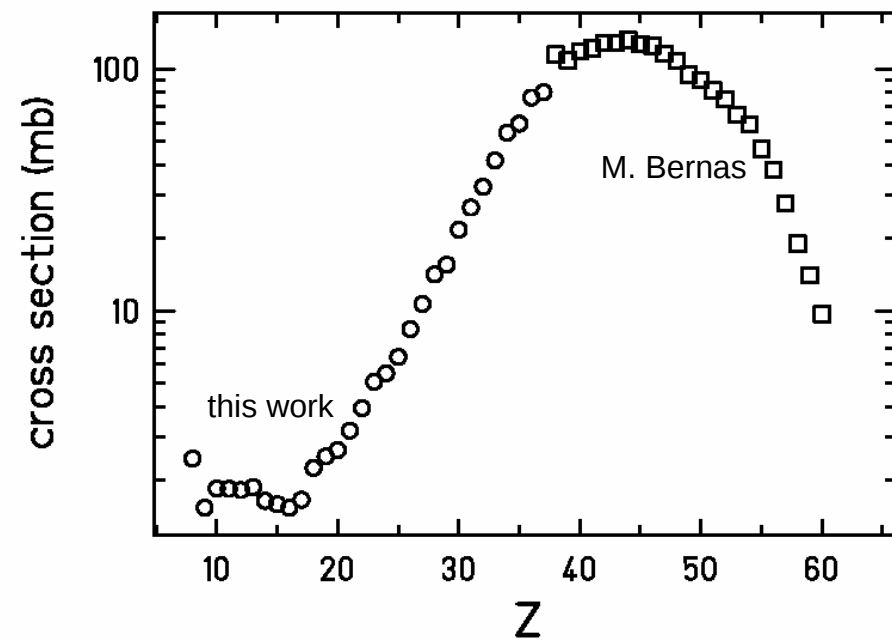
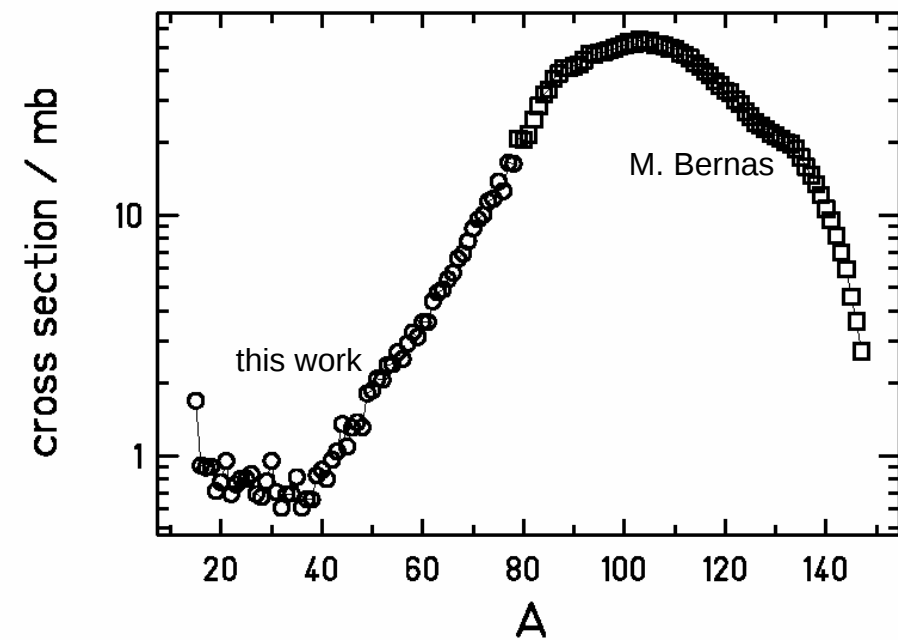
PART II:

RESULTS

CROSS SECTIONS OF THE LIGHT RESIDUES

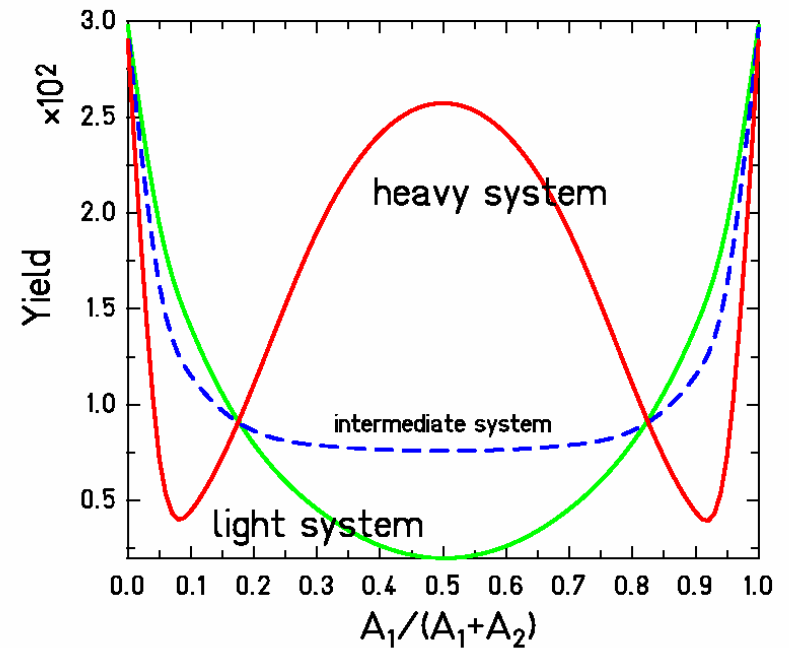
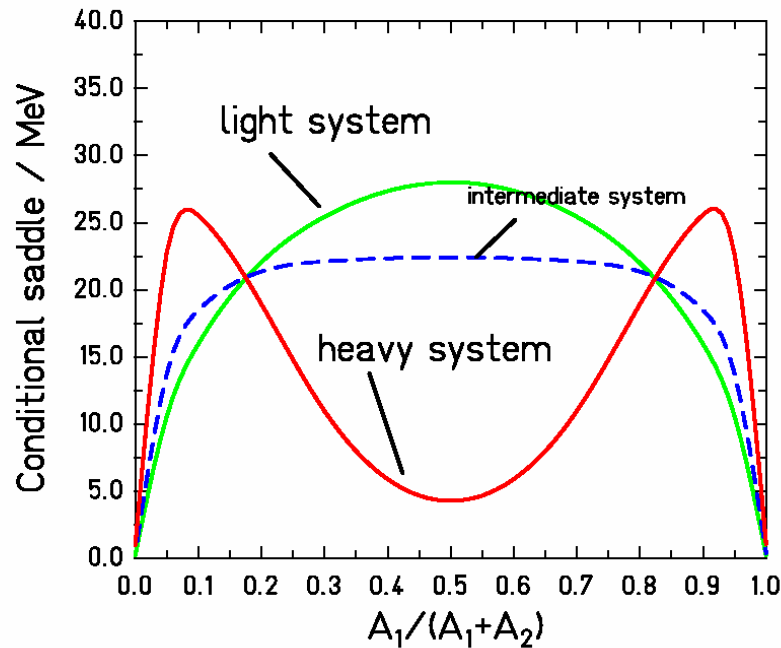


CROSS SECTIONS: MASS AND CHARGE DISTRIBUTIONS

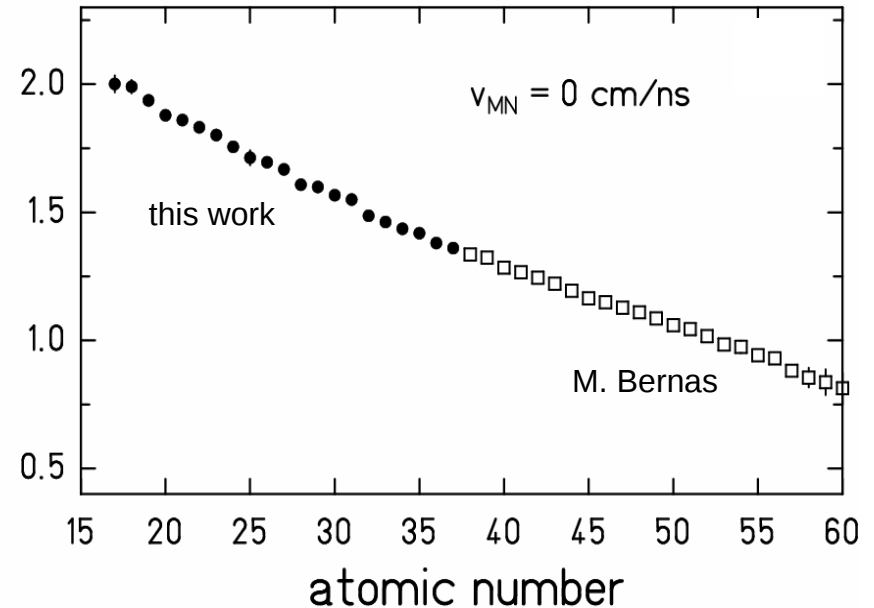
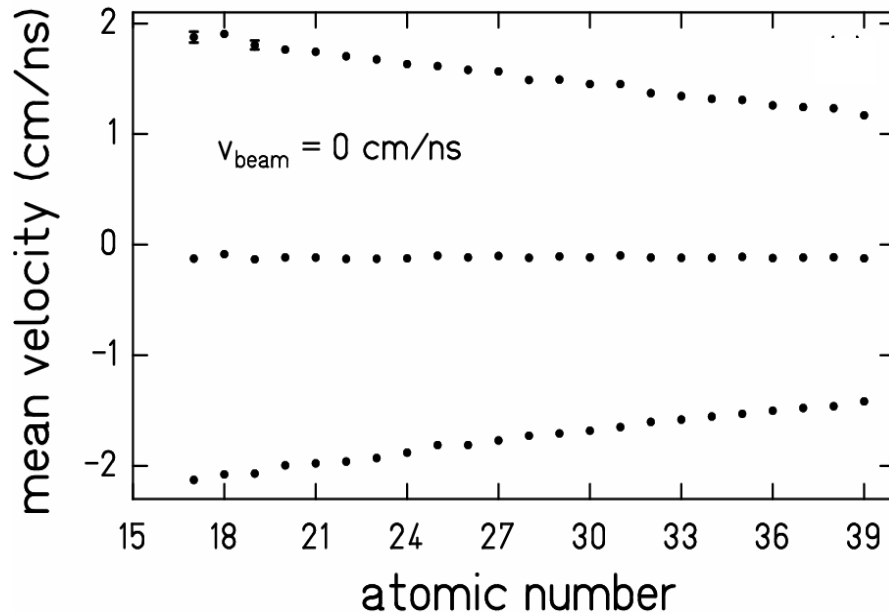


TRANSITION FROM FISSION TO EVAPORATION

Mass distribution: the statistical transition-state model



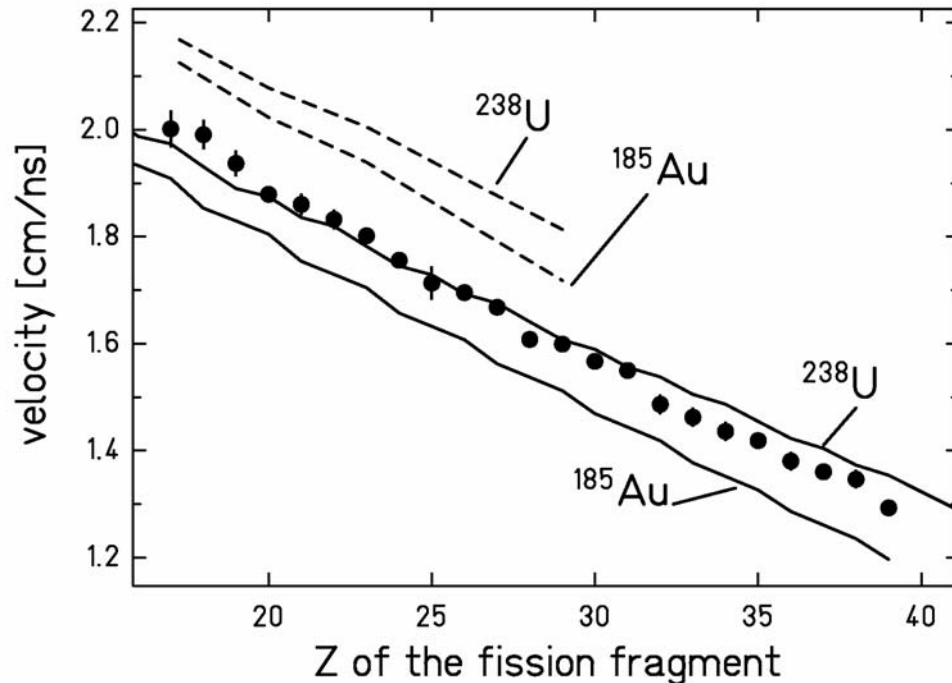
VELOCITY OF THE RESIDUES



We can interpret the counts in the two wings as very asymmetric fission products in proton-induced reactions on ^{238}U

TRANSITION FROM FISSION TO EVAPORATION

Velocity: from scission-point model towards
asymmetric decay from undeformed nucleus



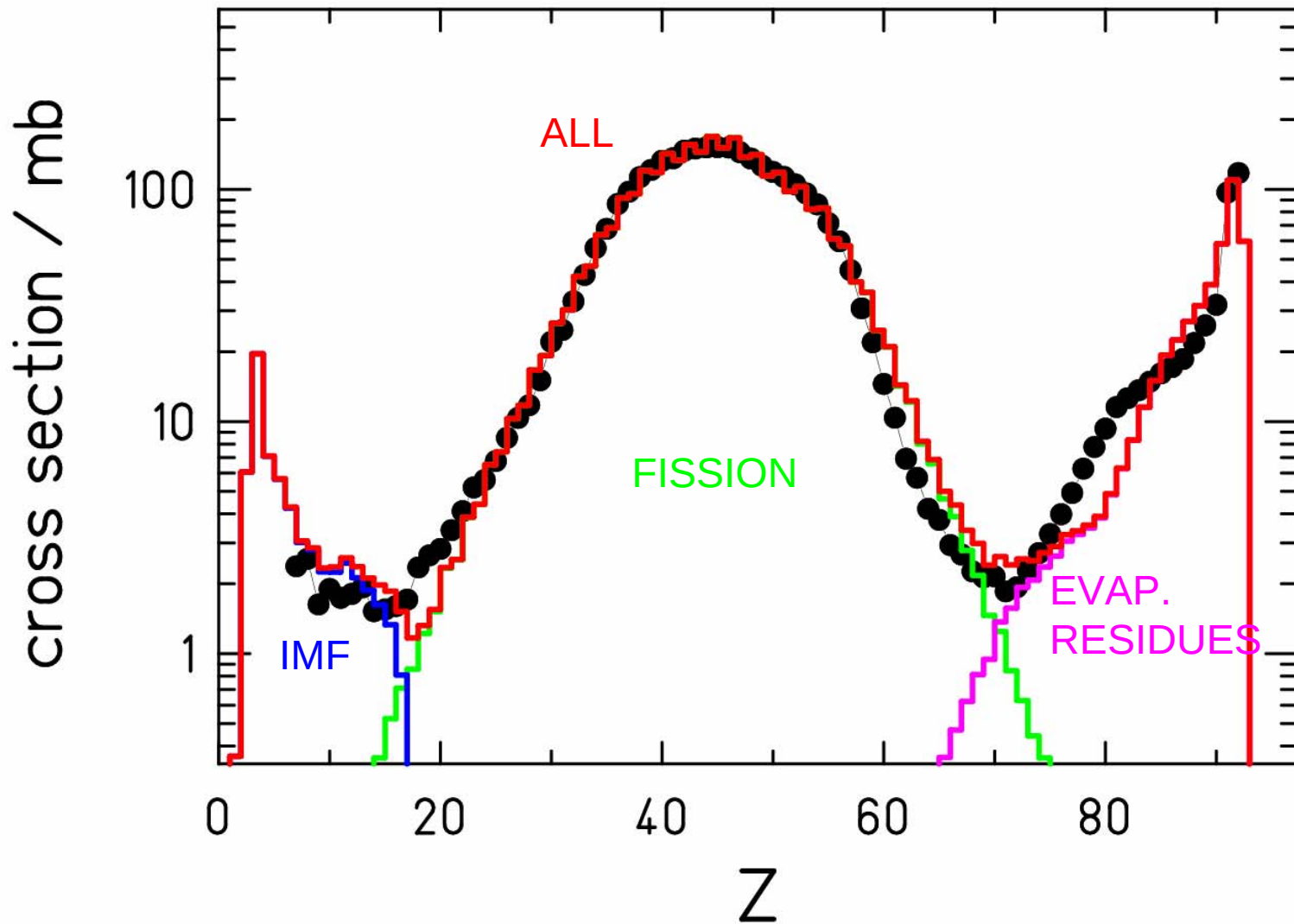
- Mean velocities in the frame of the fissioning nucleus

^{238}U , ^{185}Au = compound nuclei

— scission-point model
(deformed nuclei)

- - - Nucleus-nucleus fusion
approach (undeformed nuclei)

RESULT FOR 1 GeV p on ^{238}U



CONCLUSIONS: LIGHT NUCLIDES FROM 1 A·GeV ^{238}U + p

The experimental data:

- Production cross-sections for every isotope + velocity distributions

The data analysis:

- Details available in my PhD Thesis

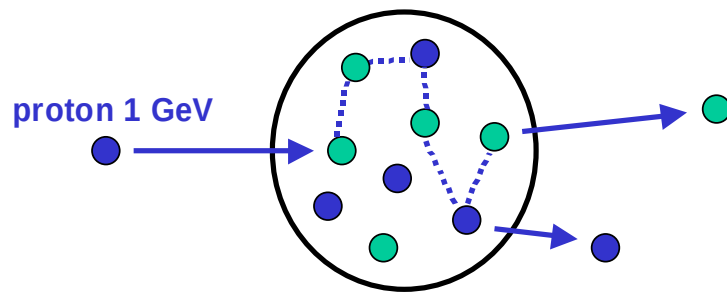
The physics:

- Yields and velocities of light nuclides from 1 A·GeV ^{238}U + p indicate a transition from fission to evaporation
- The experimental data could be successfully reproduced by a statistical model, combining a fission approach with the evaporation of intermediate-mass fragments
- No indications for a multifragmentation-type of decay have been observed

The paper:

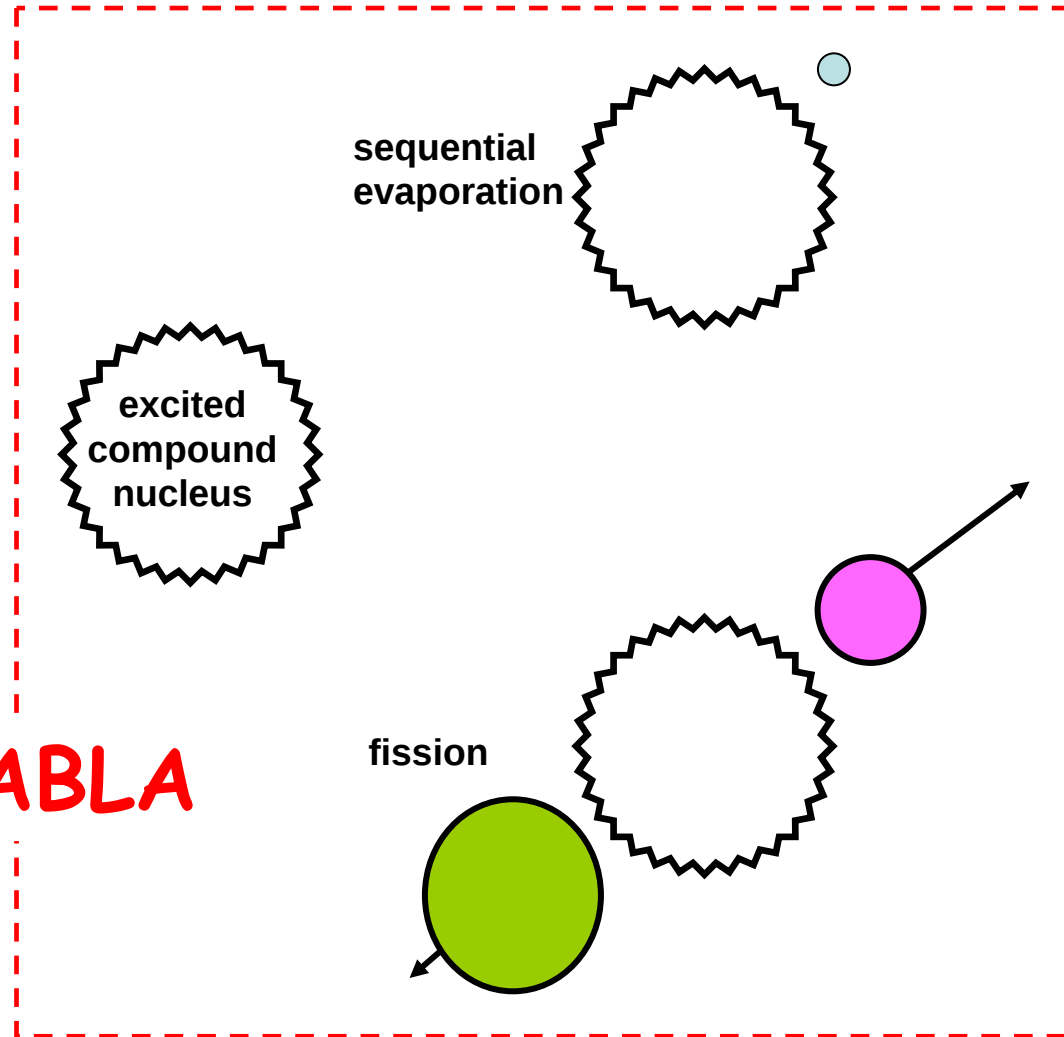
- accepted in Phys. Rev. C.

OUR STATISTICAL MODEL



Intra-nuclear Cascade

ABLA



THE FISSION MODEL: THE FISSION PROBABILITIES

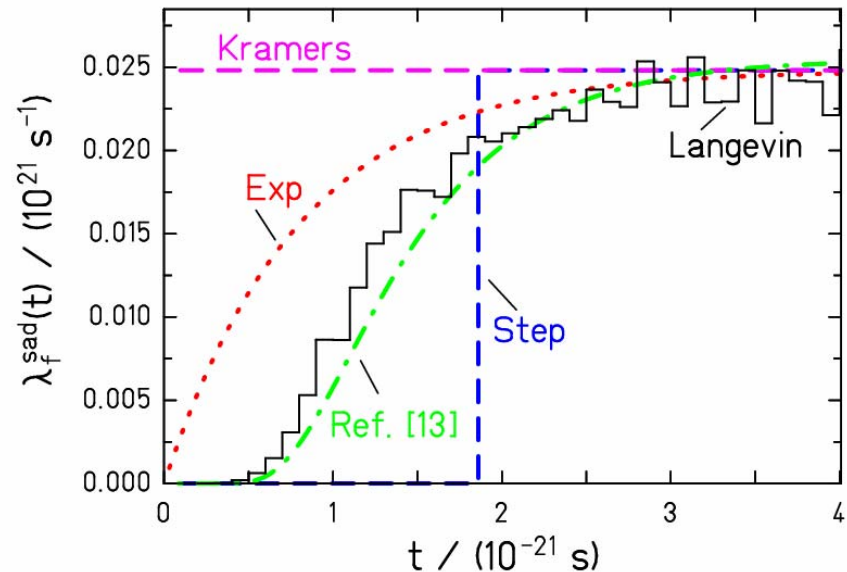
Fission decay
width: transition-
state method of
Bohr and Wheeler

$$\Gamma_f^{BW} \approx T_f \frac{\rho_f(E - B_f)}{\rho_C(E)}$$

level density of the
compound nucleus

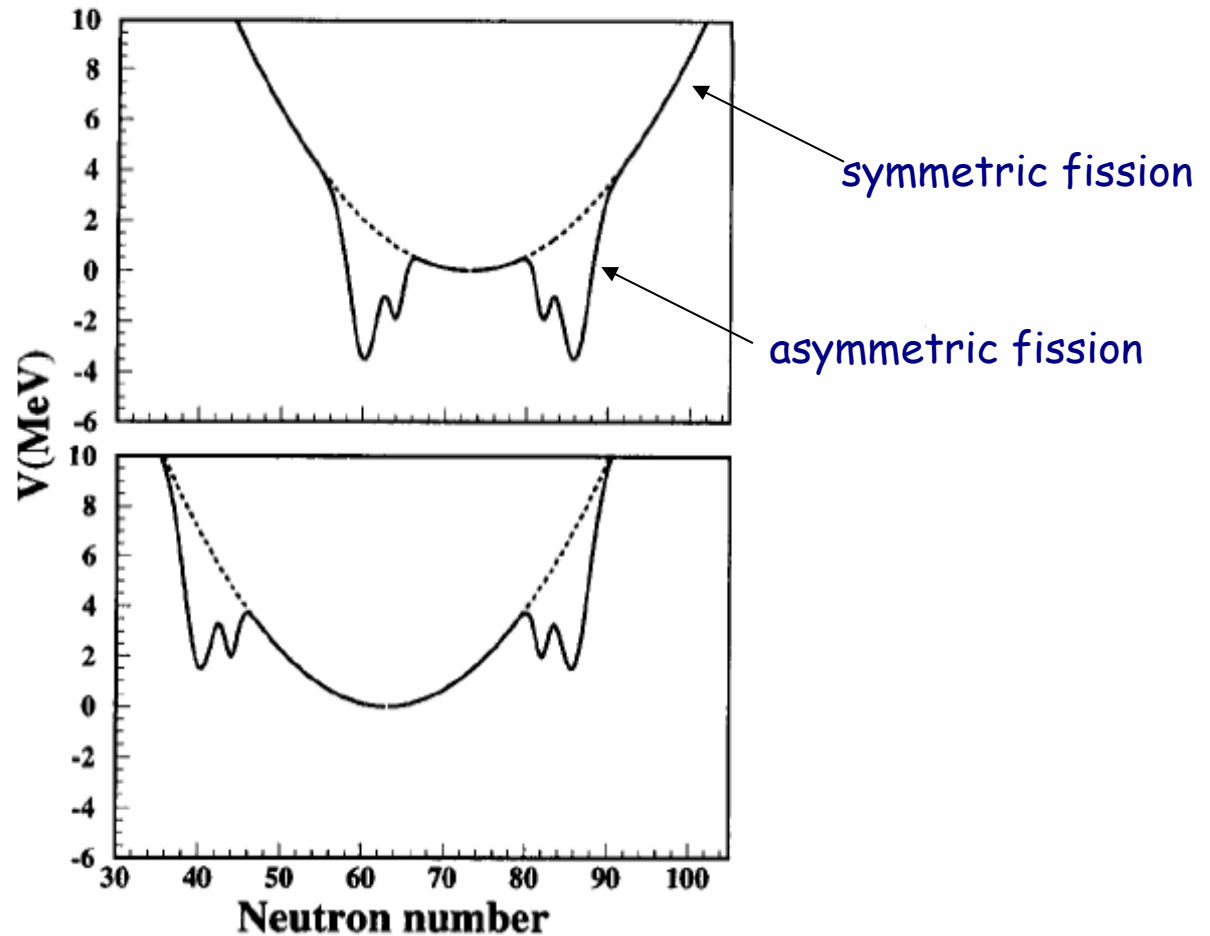
level density of the
transition states in
the saddle-point
configuration

The fission decay
widths depend
explicitly on time
(dynamical evolution of
the system along its
path to fission)
(influence of nuclear
viscosity)



THE FISSION MODEL: THE FISSION-FRAGMENT PROPERTIES (SEMIEMPIRICAL)

J. Benlliure et al. / Nuclear Physics A 628 (1998) 458–478



The modeling of
the potential
for very
asymmetric
fission is
missing

We use another
approach...

Fig. 1. Potential energy at the fission barrier for ^{238}U (upper part) and ^{208}Pb (lower part), as a function of mass asymmetry expressed by the neutron number of one of the preformed fragments.

THE VERY ASYMMETRIC FISSION: THE INTERMEDIATE-MASS FRAGMENT EMISSION

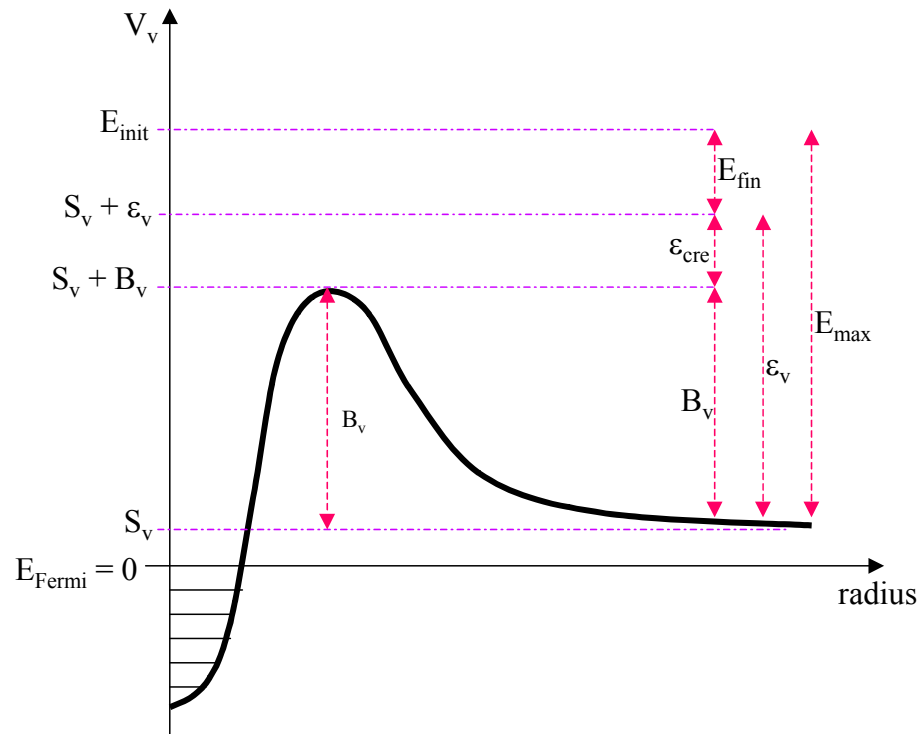
IMF decay width

$$\Gamma_{IMF} \approx \sigma_{inv} T_M^2 \frac{\rho_M(E - B_{IMF})}{\rho_C(E)}$$

level density of the
compound nucleus

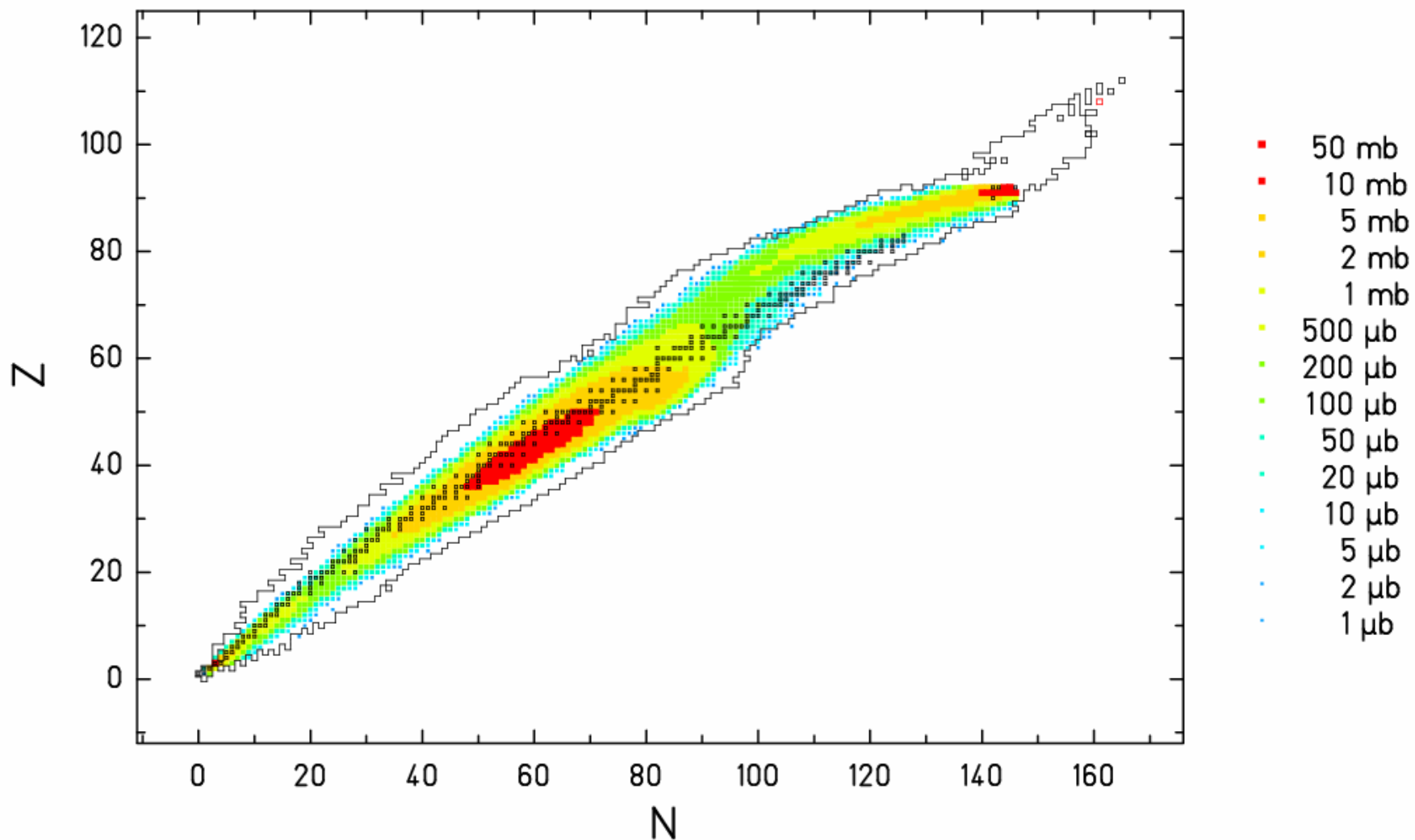
level density of
the mother
nucleus at the
barrier

The barrier is
calculated using
the fusion
nuclear
potential of
Bass



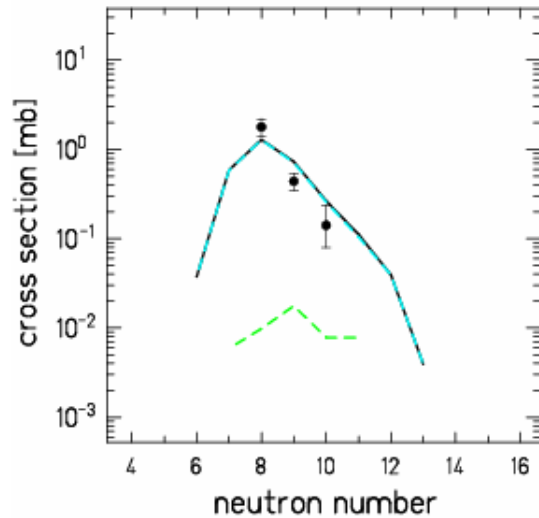
RESULT FOR 1 GeV p on ^{238}U

INCL+ABLA 1 GeV p on ^{238}U

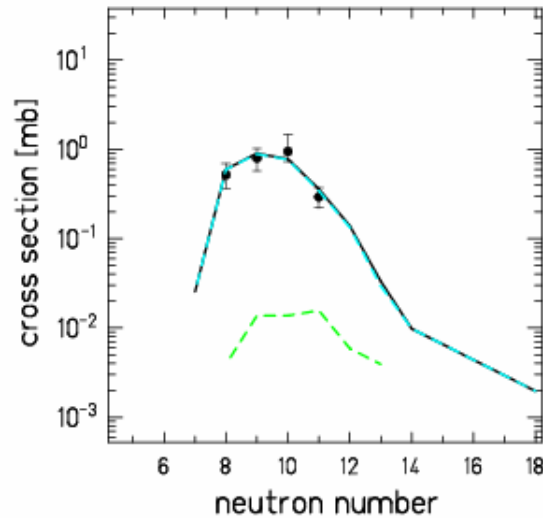


RESULT FOR 1 GeV p on ^{238}U

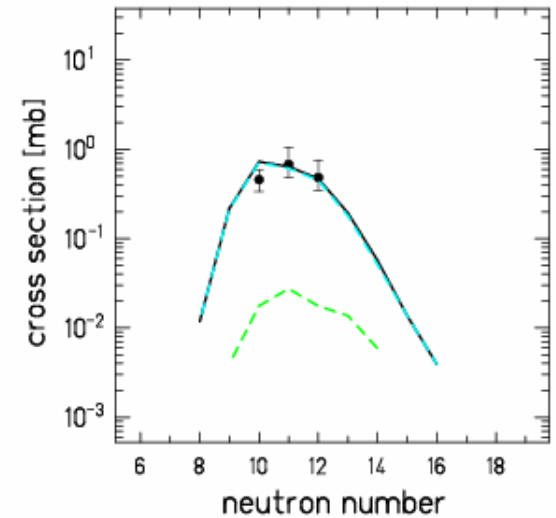
Z = 7



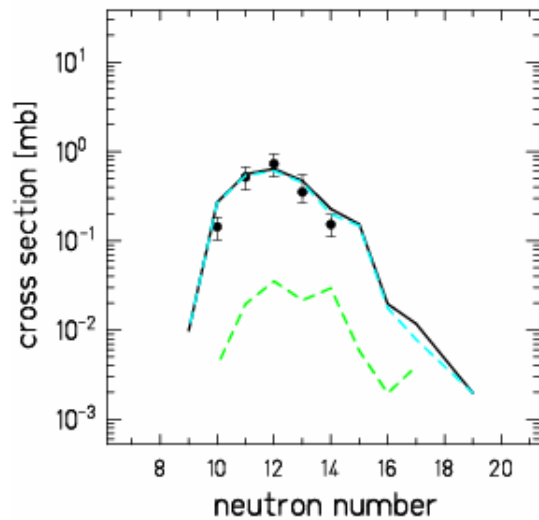
Z = 8



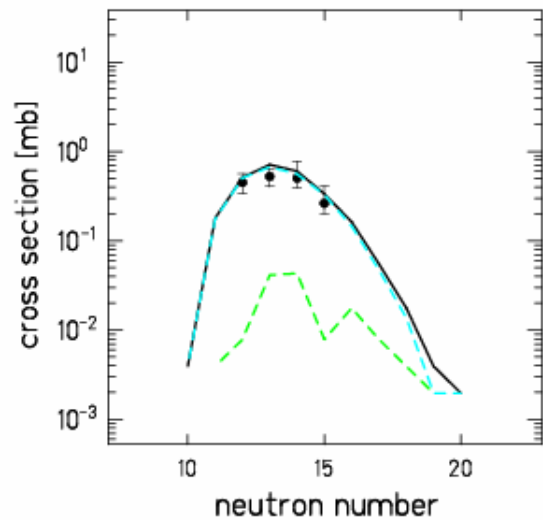
Z = 9



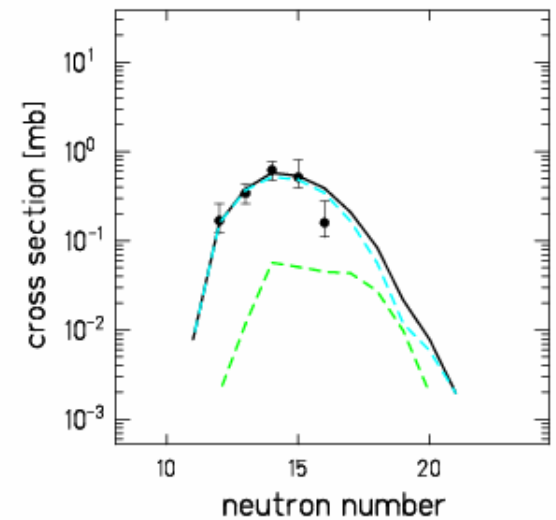
Z = 10



Z = 11

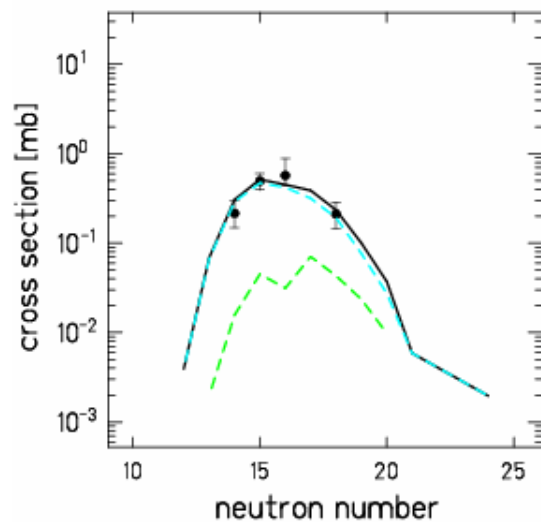


Z = 12

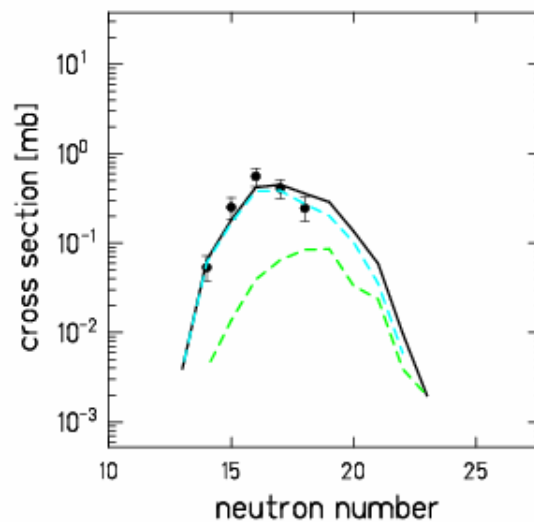


RESULT FOR 1 GeV p on ^{238}U

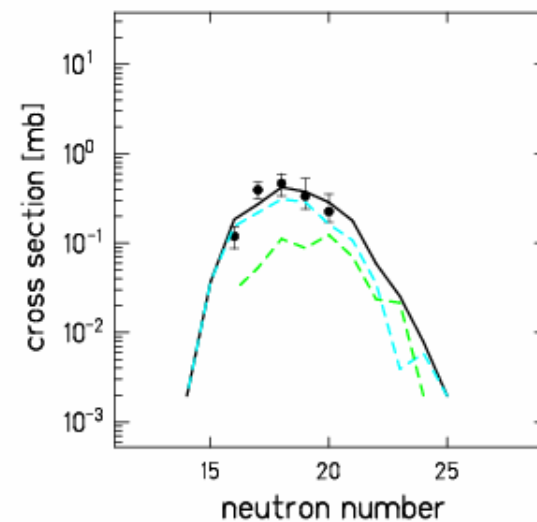
Z = 13



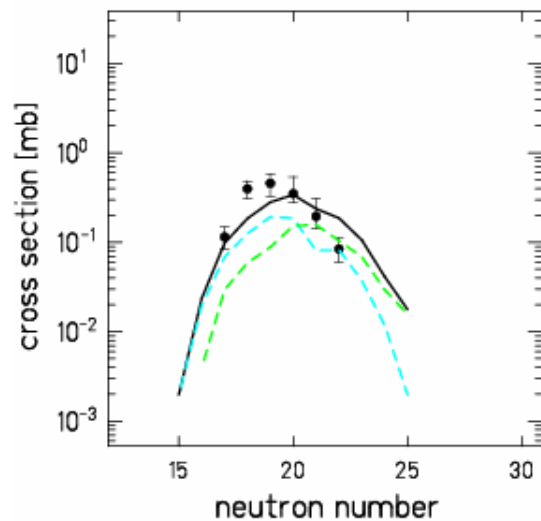
Z = 14



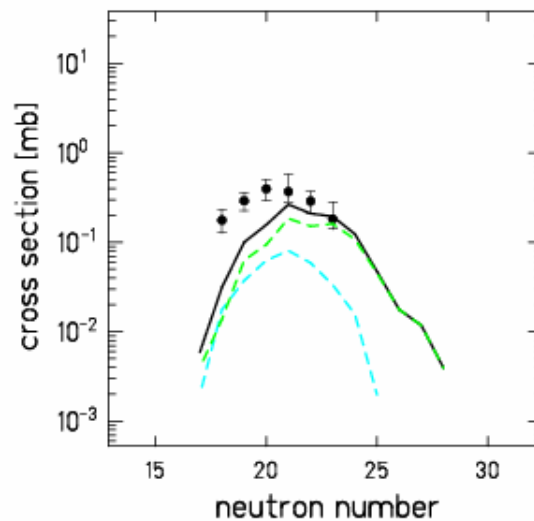
Z = 15



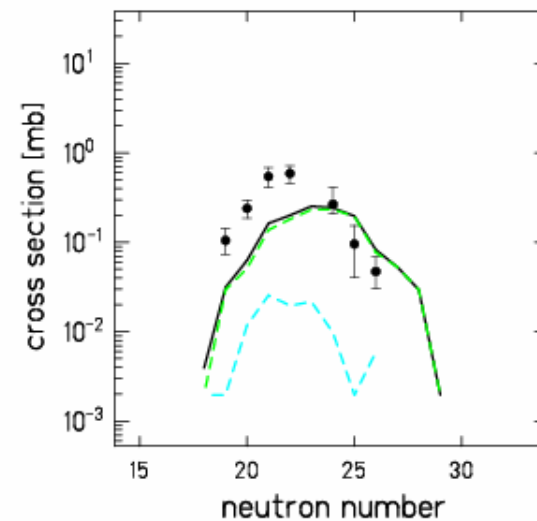
Z = 16



Z = 17

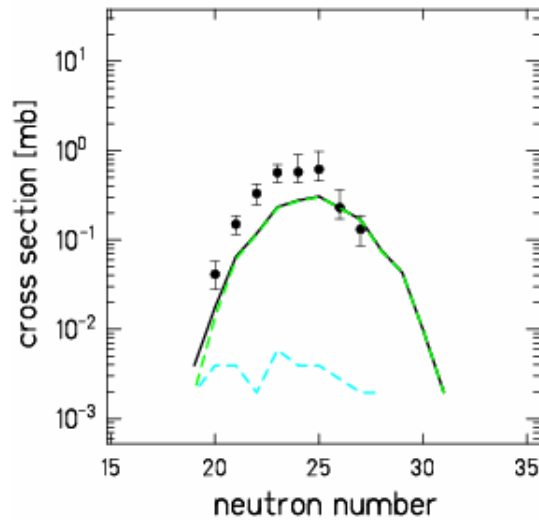


Z = 18

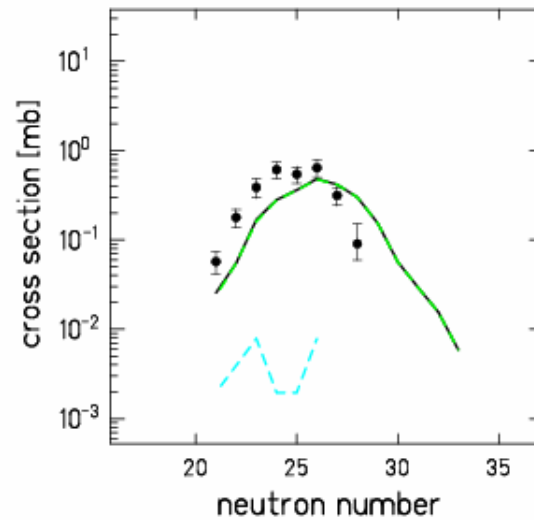


RESULT FOR 1 GeV p on ^{238}U

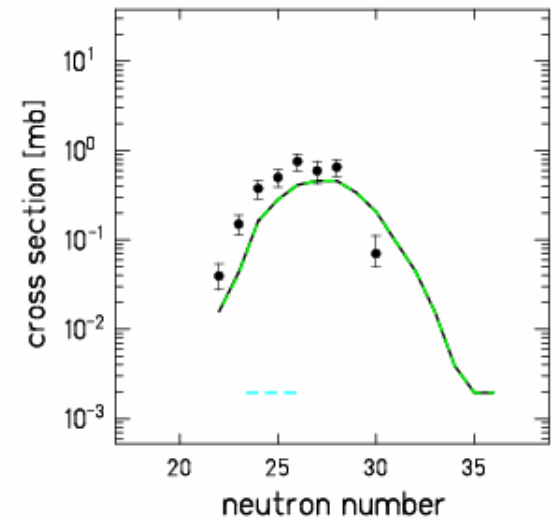
Z = 19



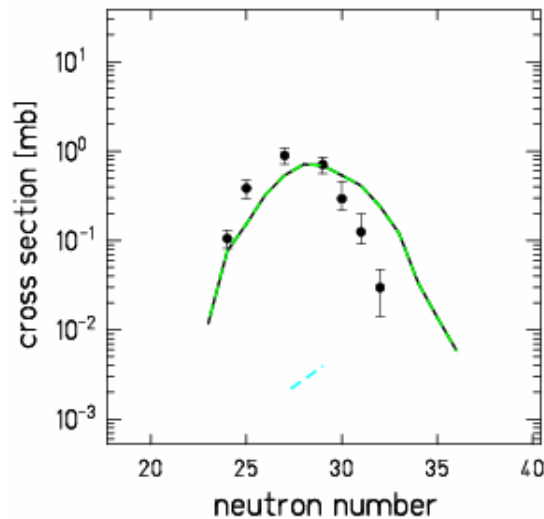
Z = 20



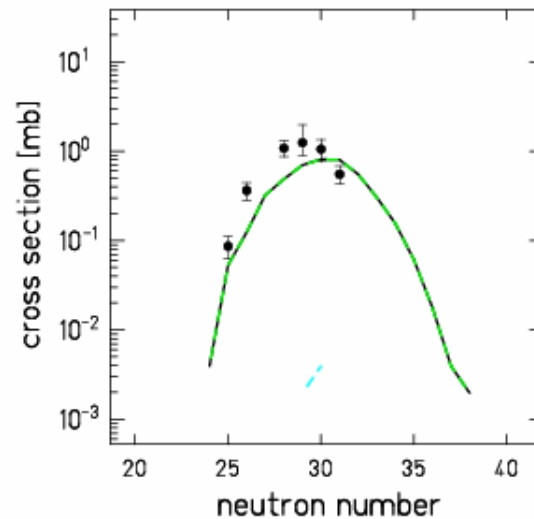
Z = 21



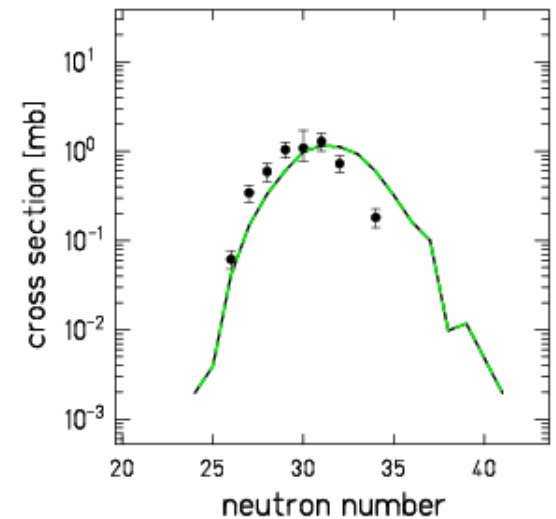
Z = 22



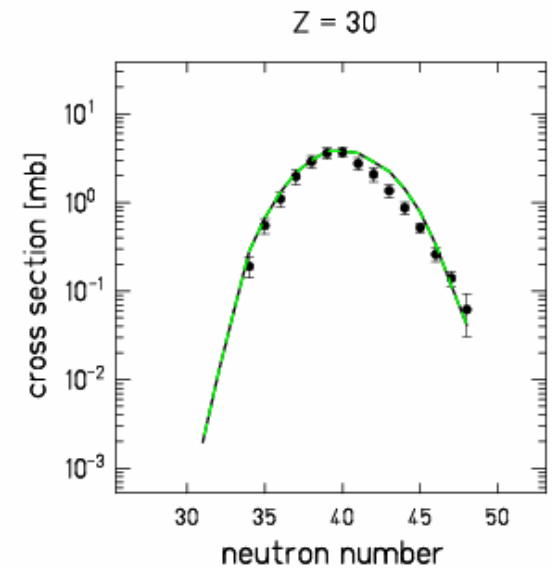
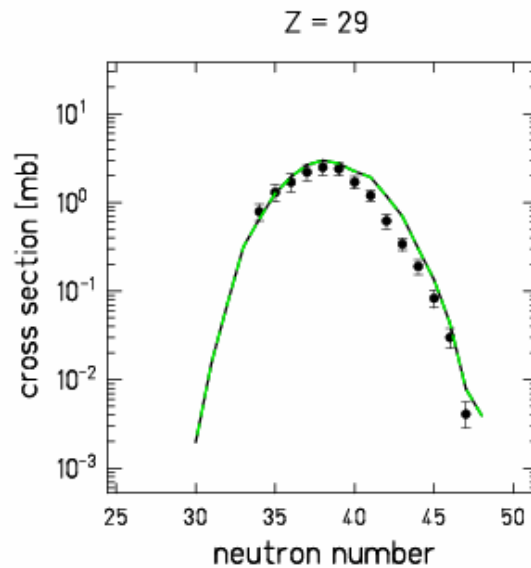
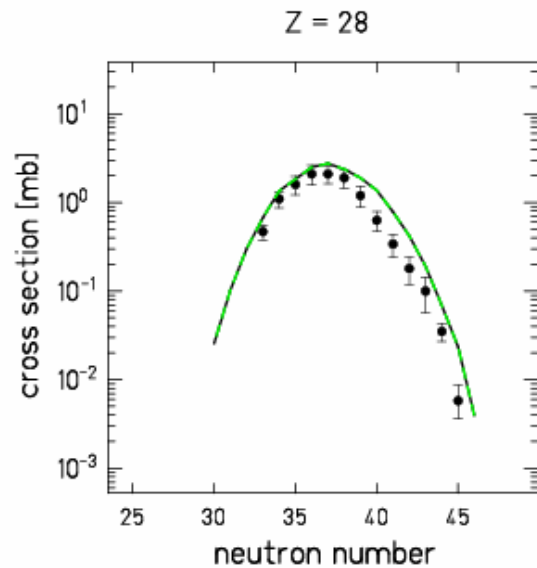
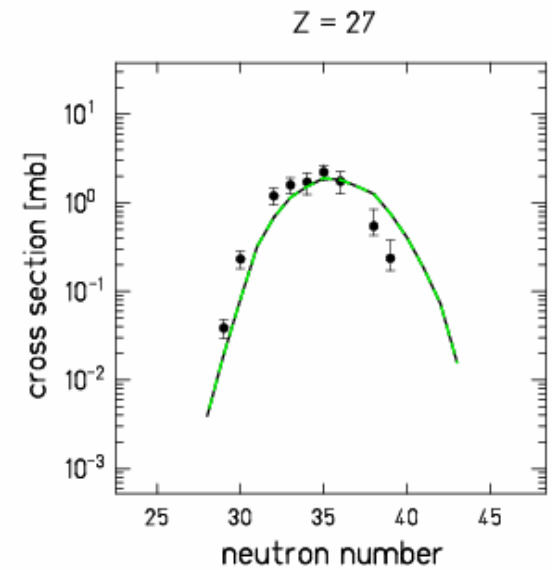
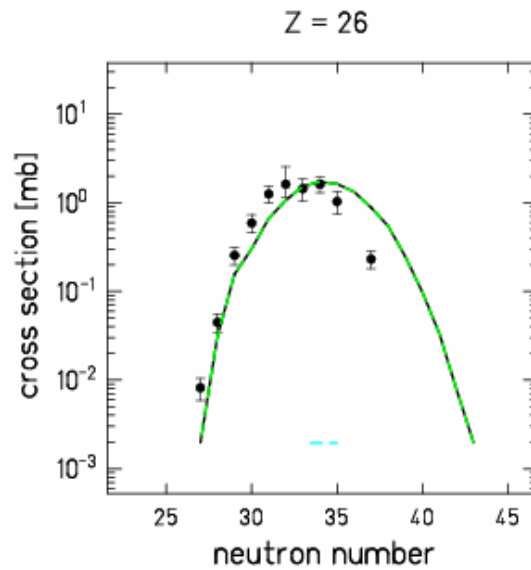
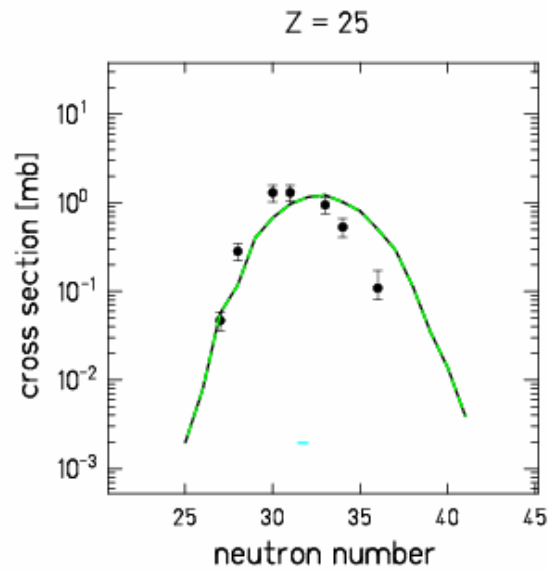
Z = 23



Z = 24

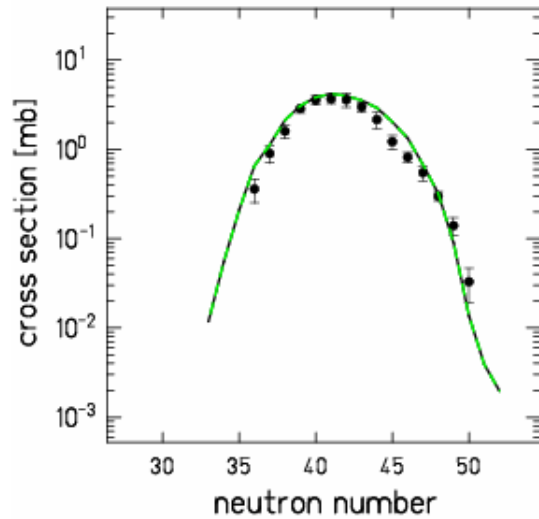


RESULT FOR 1 GeV p on ^{238}U

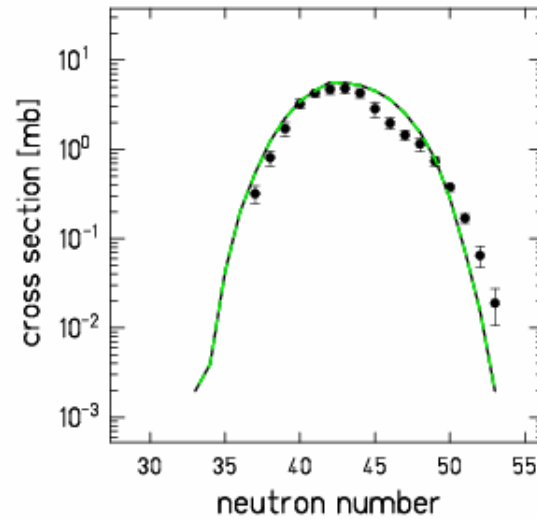


RESULT FOR 1 GeV p on ^{238}U

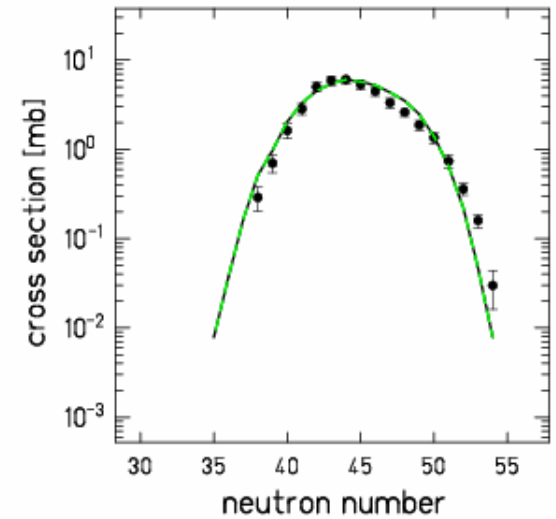
Z = 31



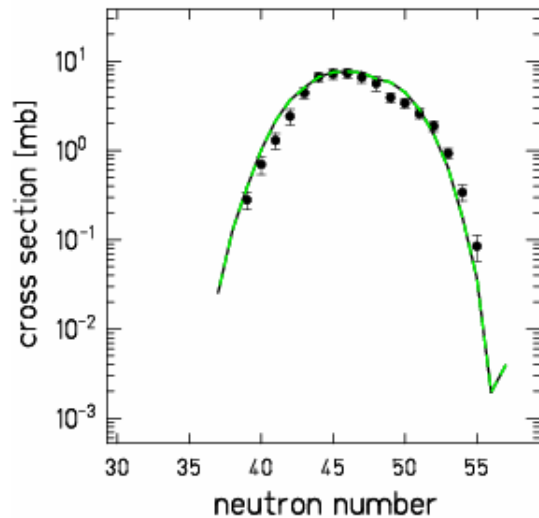
Z = 32



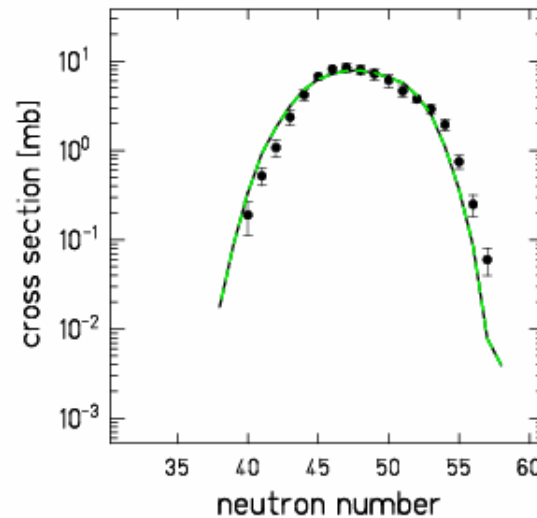
Z = 33



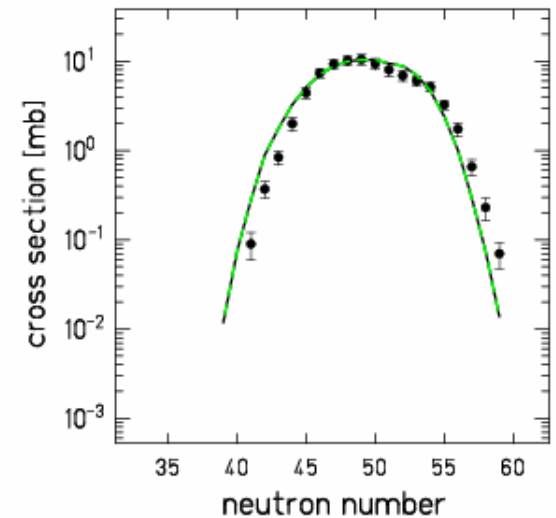
Z = 34



Z = 35

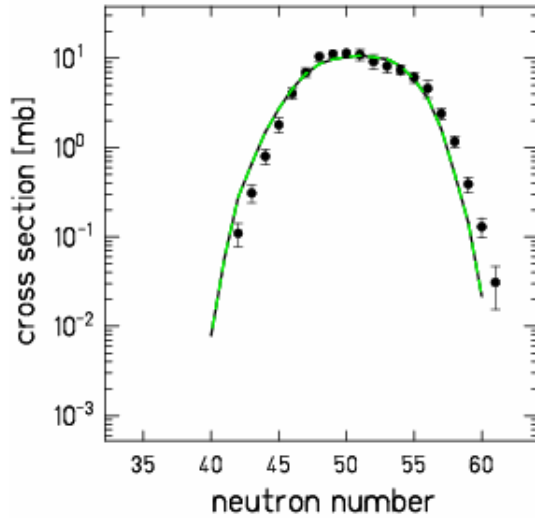


Z = 36

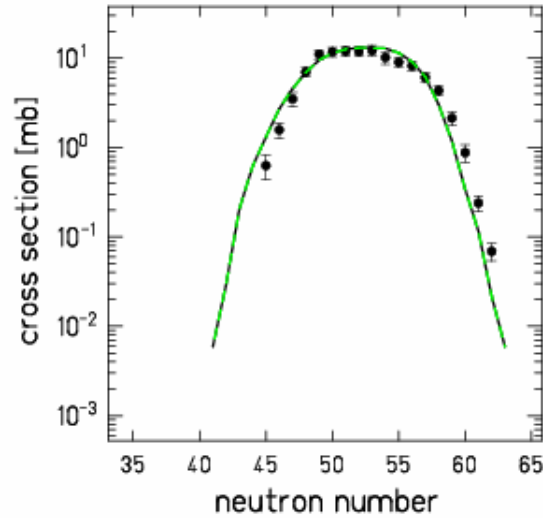


RESULT FOR 1 GeV p on ^{238}U

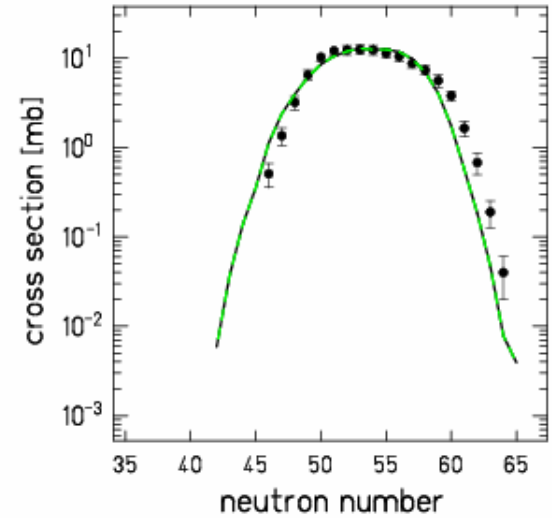
Z = 37



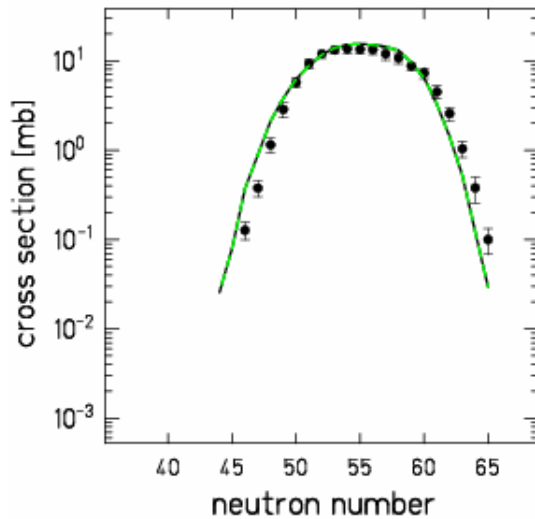
Z = 38



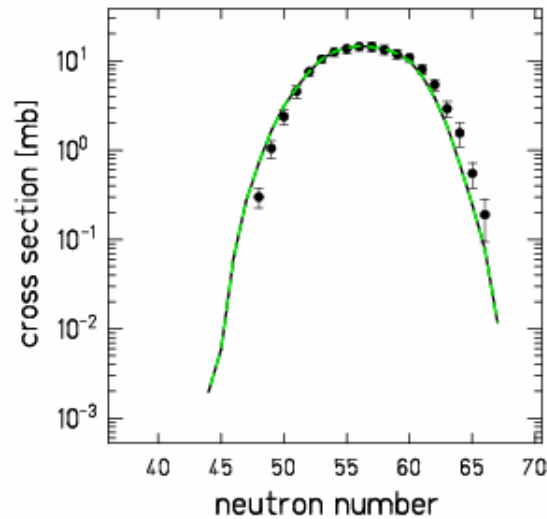
Z = 39



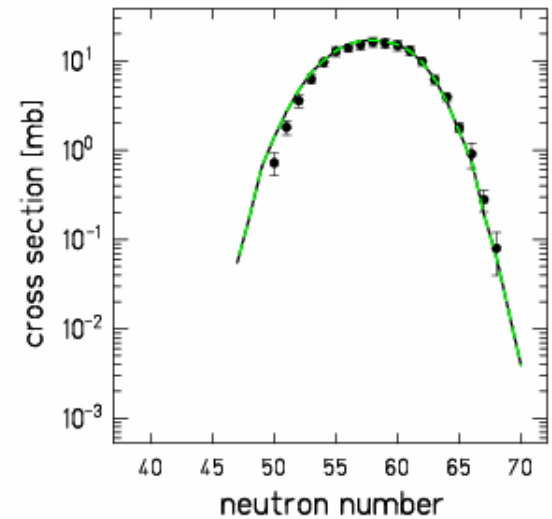
Z = 40



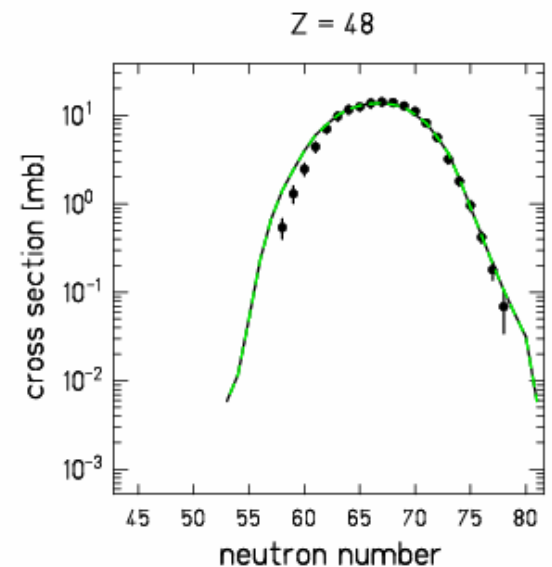
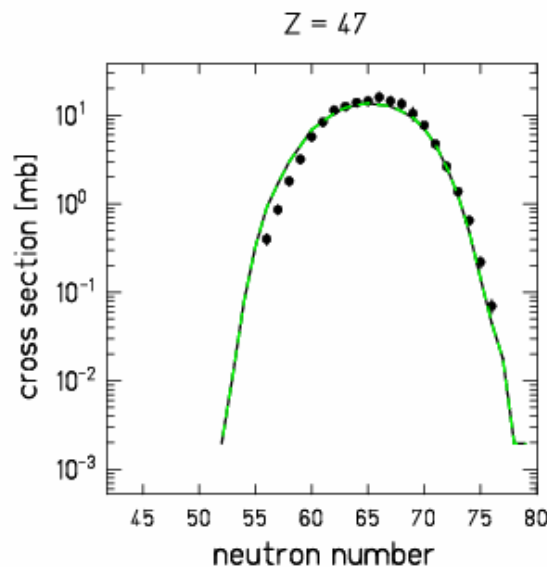
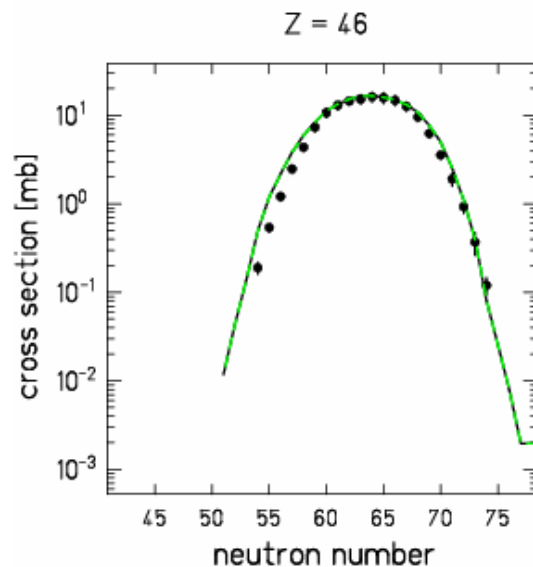
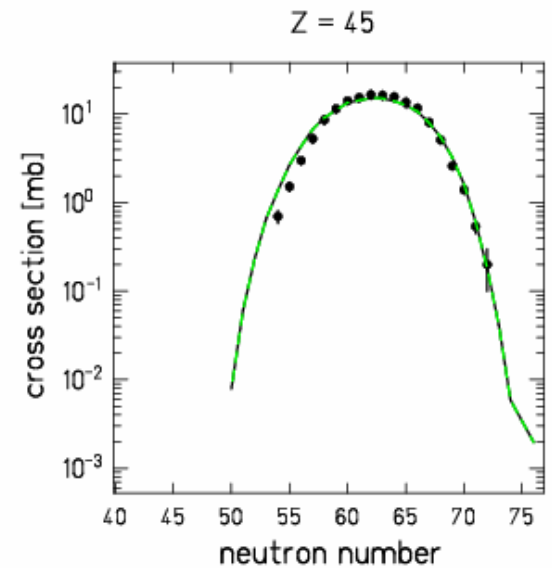
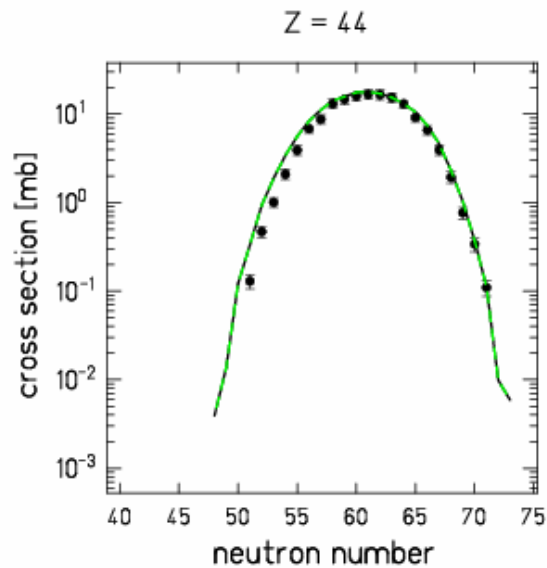
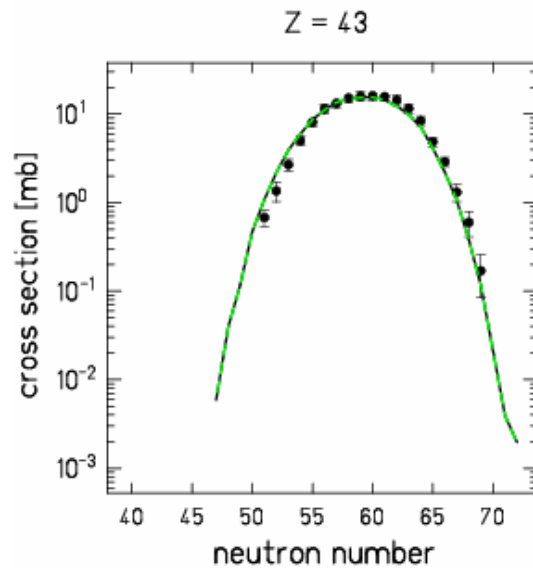
Z = 41



Z = 42

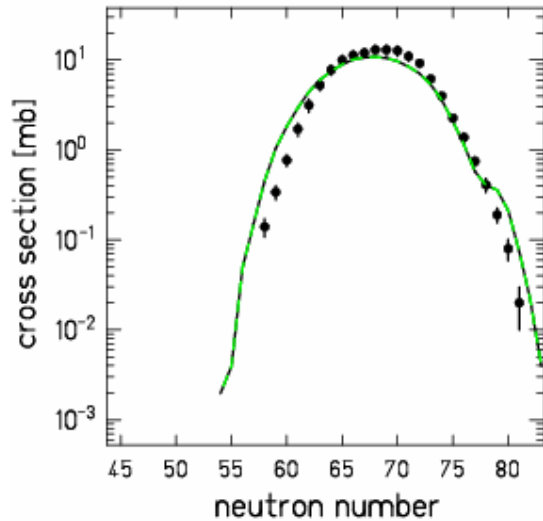


RESULT FOR 1 GeV p on ^{238}U

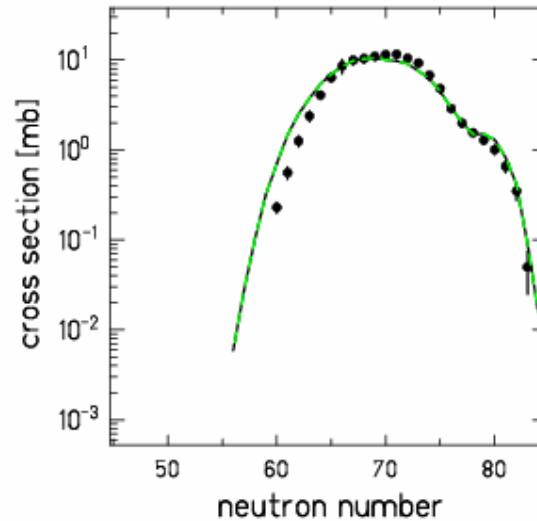


RESULT FOR 1 GeV p on ^{238}U

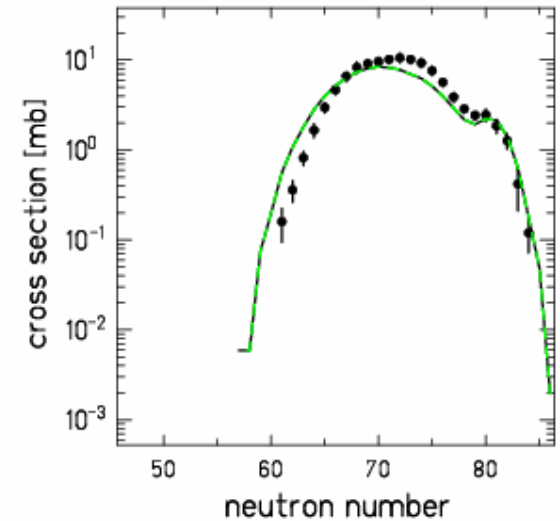
Z = 49



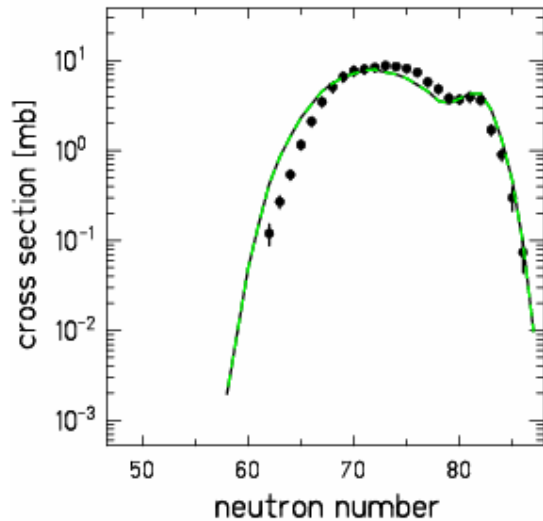
Z = 50



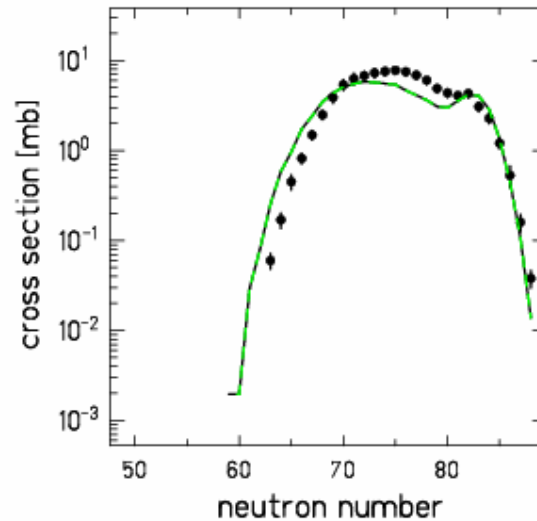
Z = 51



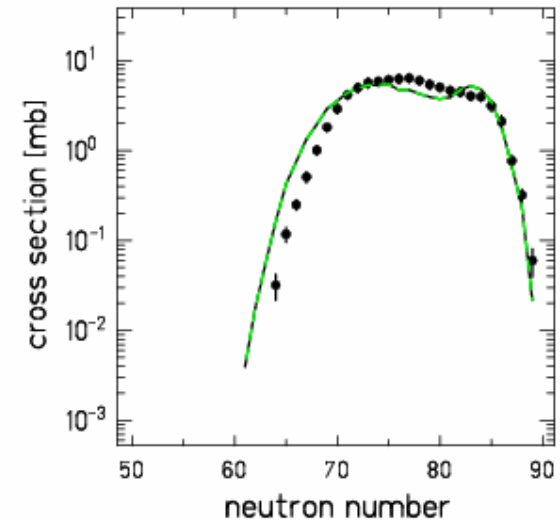
Z = 52



Z = 53

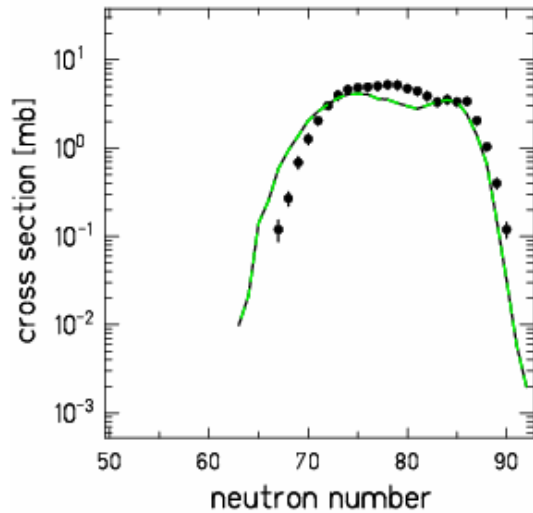


Z = 54

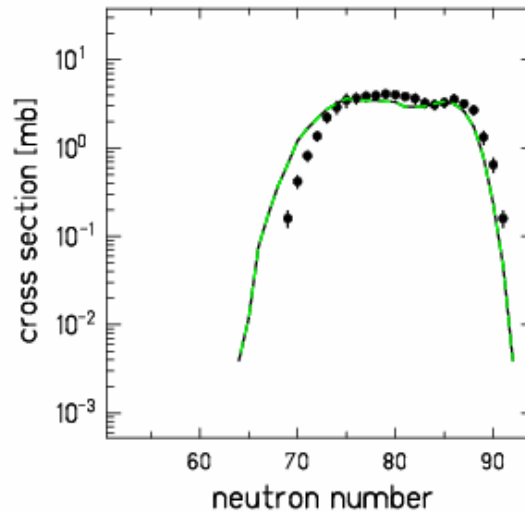


RESULT FOR 1 GeV p on ^{238}U

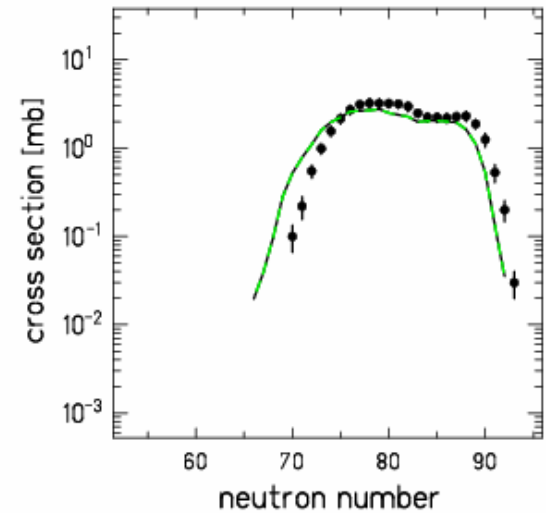
Z = 55



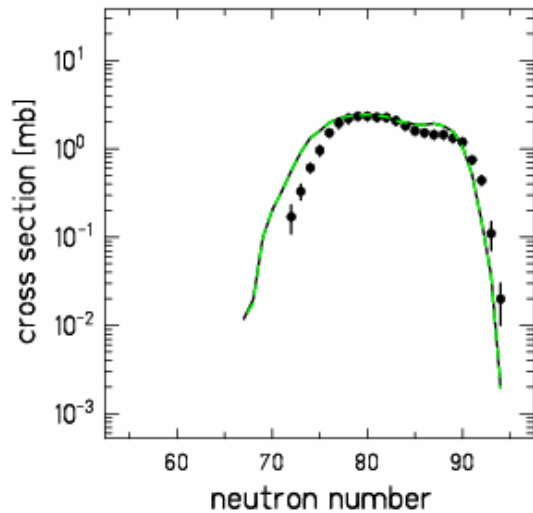
Z = 56



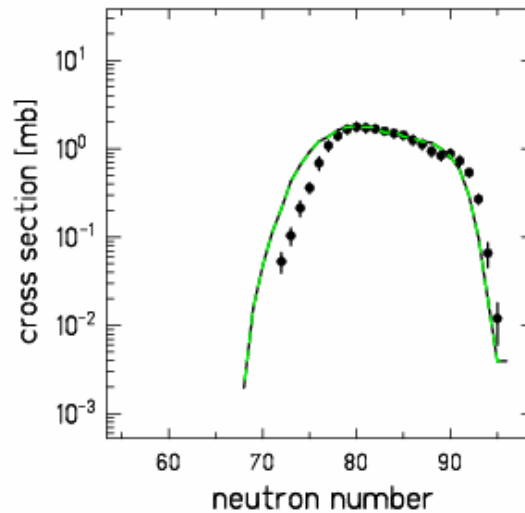
Z = 57



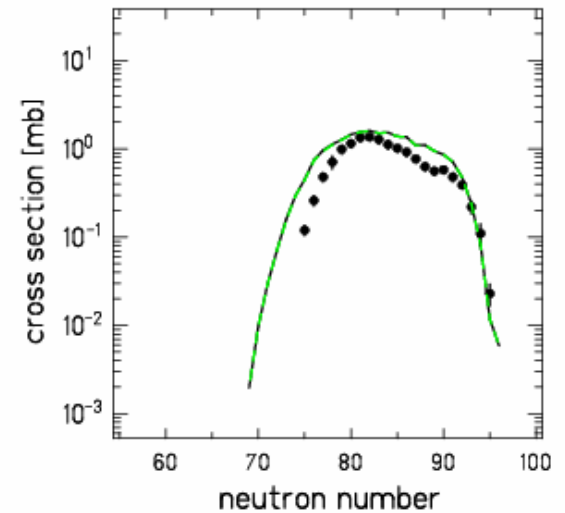
Z = 58



Z = 59

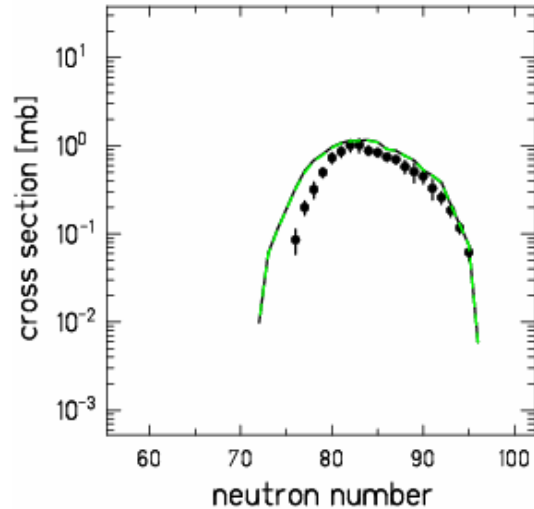


Z = 60

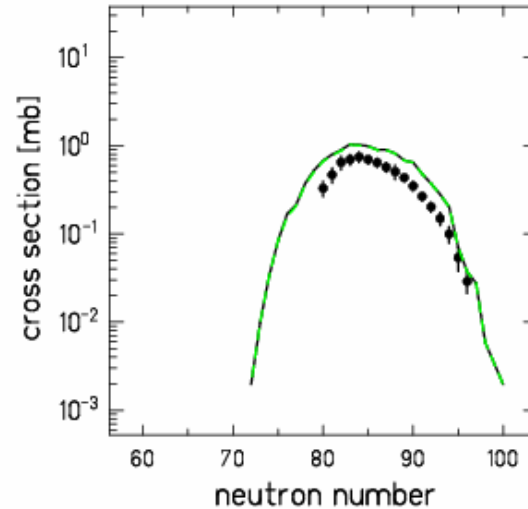


RESULT FOR 1 GeV p on ^{238}U

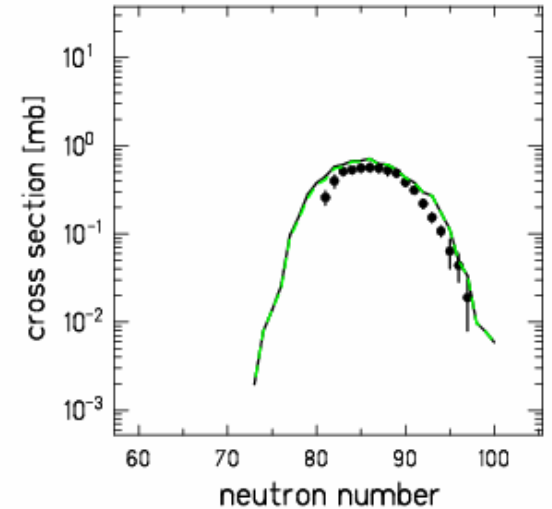
Z = 61



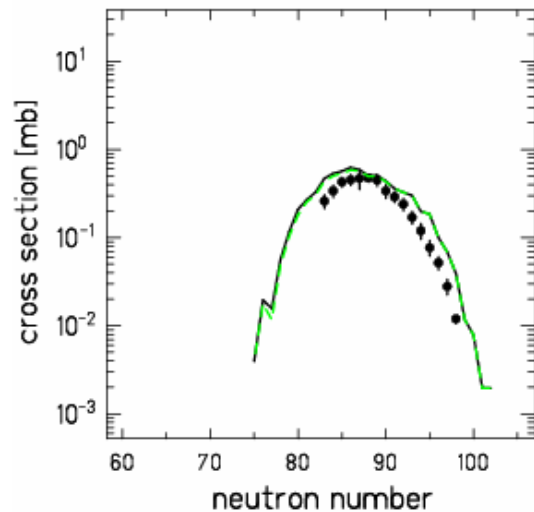
Z = 62



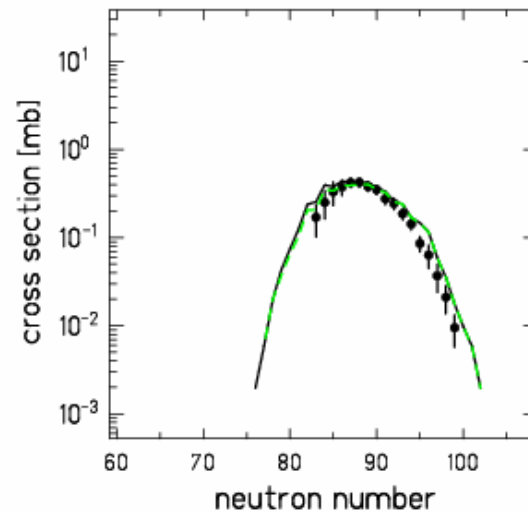
Z = 63



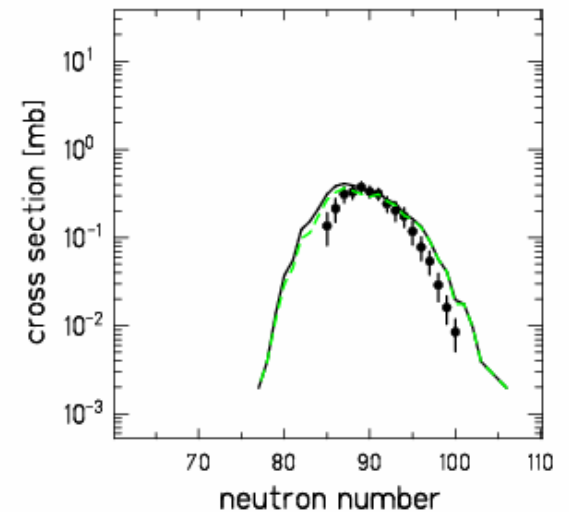
Z = 64



Z = 65

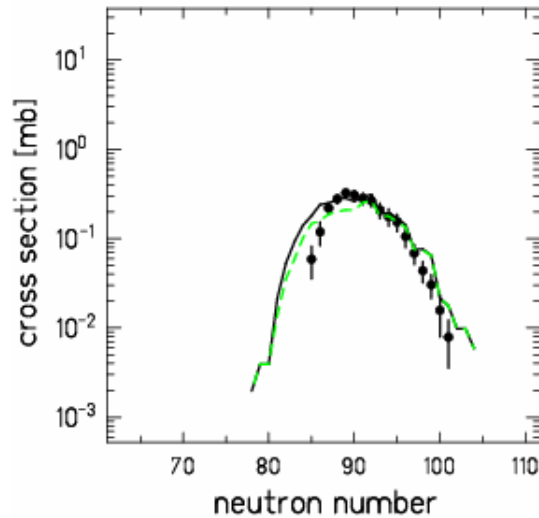


Z = 66

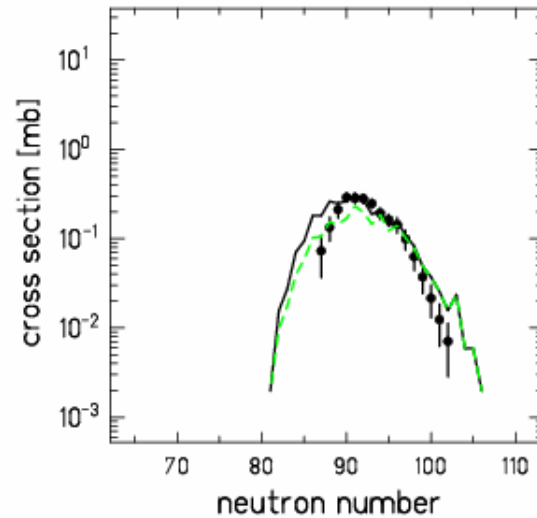


RESULT FOR 1 GeV p on ^{238}U

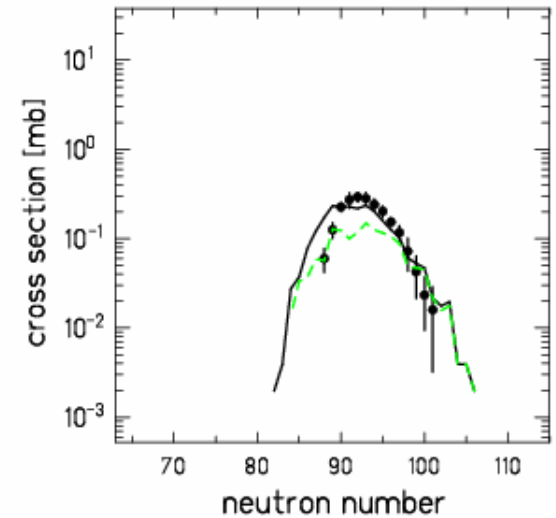
Z = 67



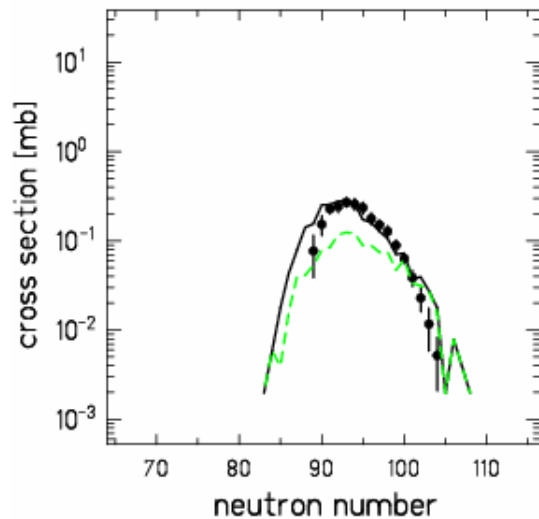
Z = 68



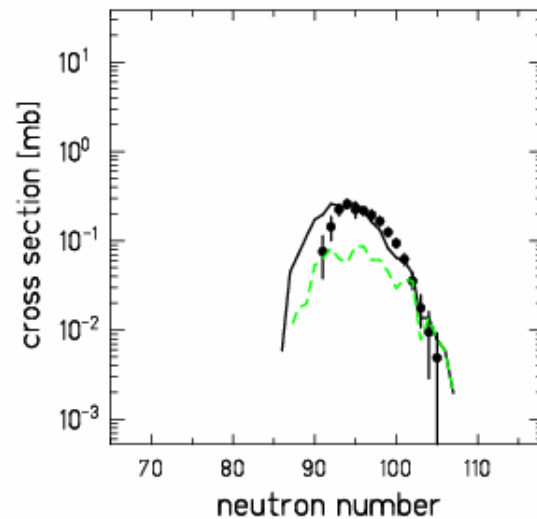
Z = 69



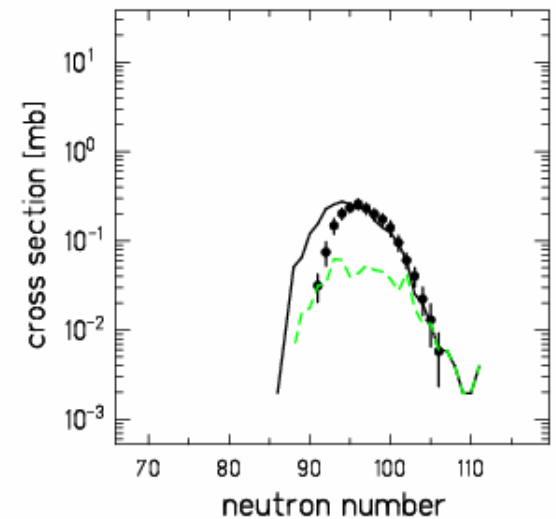
Z = 70



Z = 71



Z = 72



ADDITIONAL EFFECTS

- 1 — Thermal expansion of the nucleus
- 2 — Pre-formation factor
- 3 — Surface effects on the level density
- 4 — Deformation of the nucleus

For light systems → Thermal instabilities (break-up)

Partial widths Γ_n and Γ_p for emission of neutrons and protons.

$$\Gamma_n = \frac{2mR^2 g}{2\pi\rho(E - E_r^{gs})} \int_0^{E-B_n} \varepsilon \rho(E - B_n - \varepsilon) d\varepsilon$$

$$\Gamma_p = \frac{2mR^2 g}{2\pi\rho(E - E_r^{gs})} \int_{\varepsilon_c}^{E-B_p} \varepsilon \left(1 - \frac{\varepsilon_c}{\varepsilon}\right) \rho(E - B_p - \varepsilon) d\varepsilon$$

(Approximation without considering tunneling.)

Γ_p is reduced by the Coulomb barrier ε_c .

Modelling the Width in A and N/Z of Fission-Product Isotopic Distributions

Approximated parabolic potential

$$U(\eta) = C_{\eta} \cdot (\eta - \eta_o)^2$$

Statistical population:

$$Y(\eta) \propto \exp\{2\sqrt{a(U_0 - U(\eta))}\} \quad \rightarrow$$

$$Y(\eta) \propto \exp\left\{-\frac{(\eta - \eta_o)^2}{2\sigma_{\eta}^2}\right\} \quad \text{with}$$

$$2\sigma_{\eta}^2 = \frac{T}{C_{\eta}}$$

U = **potential energy**,
 η = **either A (mass split) or N/Z (polarisation)**,
 C_{η} = **stiffness of the potential**,
 T = **nuclear temperature**.