

Fission hindrance from GDR γ -rays emission - Overview

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- GDR clock -

- * based on the emission of high energy γ rays from giant dipole resonance excited in the hot fissioning nucleus prior to scission.
- * its speed depends only on the level density parameter and the GDR strength function.
- ✓ in principle no free parameters.
- ✓ no transmission factors.
- ✓ strength function follows the evolving deformation of nucleus.
- ✓ sensitivity of the γ rays - fission angular correlation to the deformation of the nucleus at the time of γ ray emission.
- ✗ fissioning nucleus γ rays and fission fragments γ rays can not be distinguished experimentally.

- Fission time scales -

* approach of Grange and Weidenmüller

1. inside the saddle:

$$\Gamma_f(t) = \Gamma_f^{\text{KRAMERS}} \left[1 - \exp\left(-\frac{t}{\tau_f}\right) \right]$$

2. at the saddle:

$$\Gamma_f^{\text{KRAMERS}} = \Gamma_f^{\text{BW}} \left(\sqrt{1 + \gamma^2} - \gamma \right); \quad \Gamma_f^{\text{BW}} = \frac{\hbar \omega_{\text{sc}}}{2\pi} \exp\left(-\frac{B_f}{T}\right)$$

3. from saddle to scission

$$\tau_{\text{ssc}} = \tau_{\text{ssc}}^0 \left(\sqrt{1 + \gamma_0^2} + \gamma_0 \right), \quad \tau_{\text{ssc}}^0 = 3 \cdot 10^{-24} \text{ s}$$

$$\gamma \text{ (nuclear friction coeff.)} = \frac{\beta}{2\omega}$$

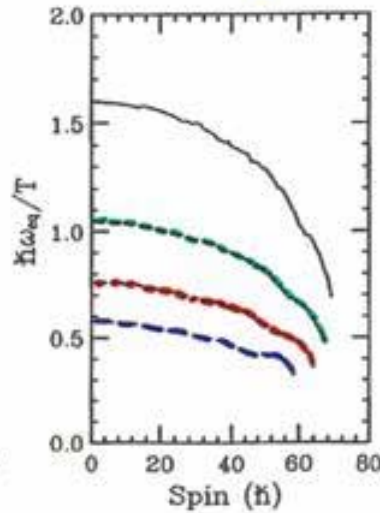
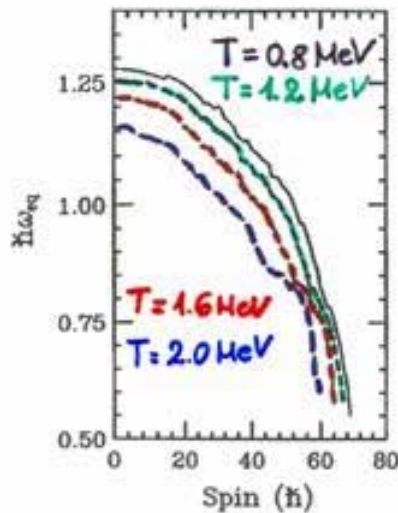
$$\beta \text{ (reduced dissipation coeff.)} = \frac{\eta}{m}$$

$$\tau_f = \begin{cases} \frac{\beta}{2\omega_f^2} \ln \frac{10B_f}{T}, & \text{overdamped motion} \\ \frac{1}{\beta} \ln \frac{10B_f}{T}, & \text{underdamped motion} \end{cases}$$

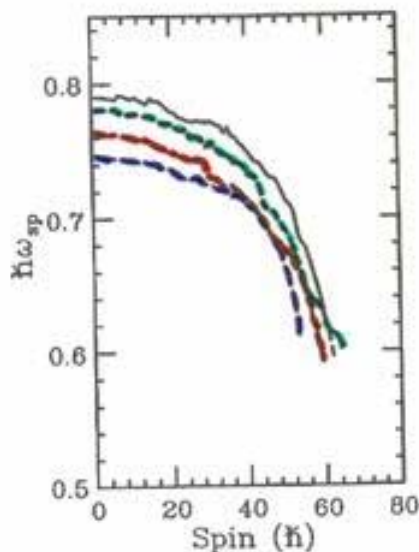
$$(\gamma_{\text{inside}}, \gamma_{\text{outside}}) \Rightarrow \tau_{\text{fis}}^{\text{TOTAL}} = \tau_f + \tau_{\text{ssc}}$$

- Cascade -

$$\Gamma_{\text{CASCADE}}^{(\gamma=0)} = \frac{T}{2J_1} \exp\left(-\frac{B_f}{T}\right) = \frac{T}{\hbar\omega_{\text{eq}}} \Gamma_f^{\text{BW}} : \hbar\omega_{\text{eq}} = T$$



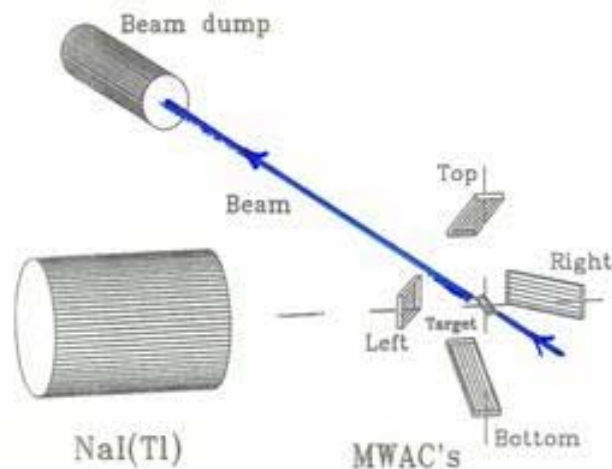
$$\gamma \leftrightarrow \beta \leftrightarrow \tilde{\gamma} : \omega_{\tilde{\gamma}} = 1 \cdot 10^{24} \text{ s}^{-1}$$



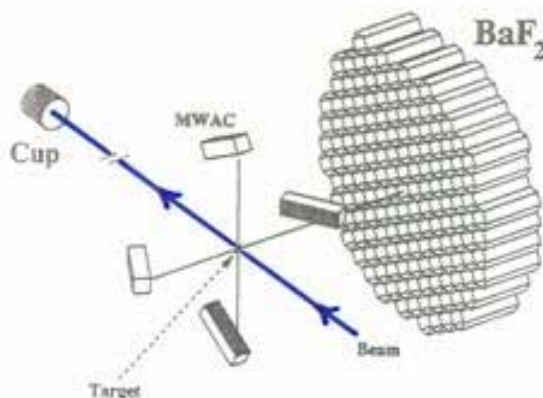
M. Thoennessen,
 Proceed. 3rd Intern. Conf.
 Dynam. Aspects of Nucl.
 Fission, Častá-Papierníka,
 Slovak Republic.

- Experimental methods -

* Diószegi J. et al., Phys. Rev. C 46 (1992) 627

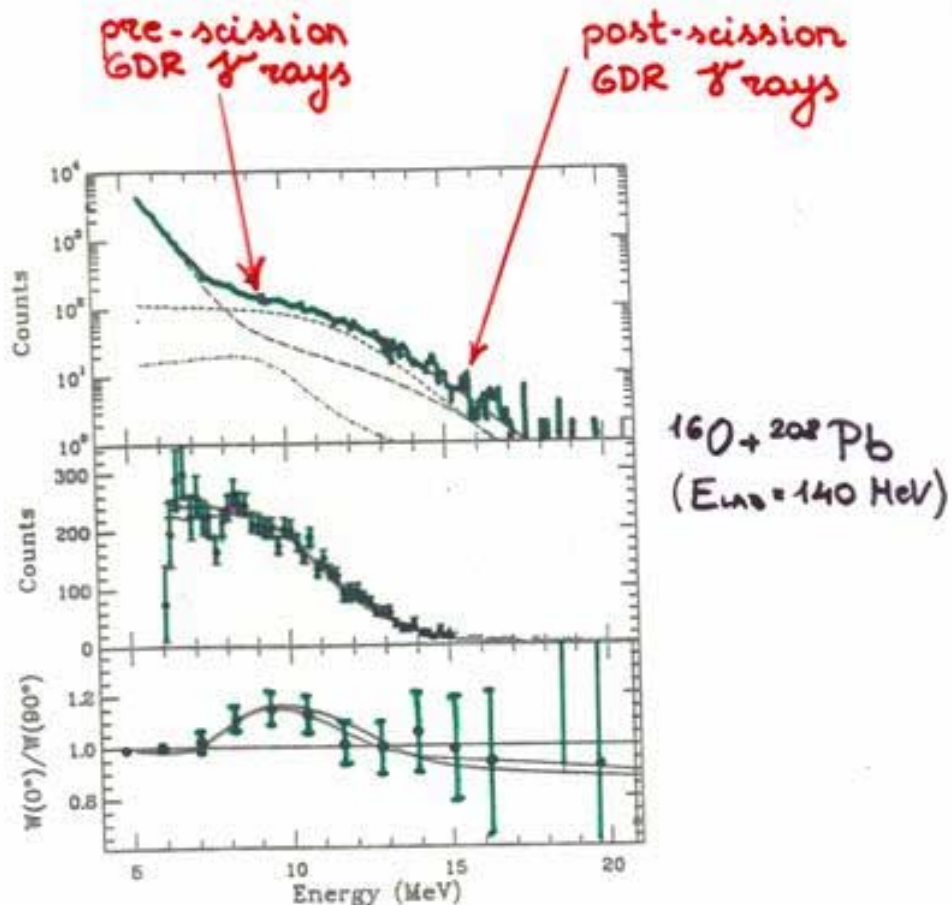


* Shaw R.P. et al., Phys. Rev. C 61 (2000) 044612



- Experimental methods -

- * medium and heavy nuclei : $E_{\text{GDR}} \approx 80 \cdot A^{-1/3} \text{ (MeV)}$.
- * simultaneous fit to the γ ray spectrum and the angular correlation can determine the strengths of the pre- and post-scission GDR decays separately.



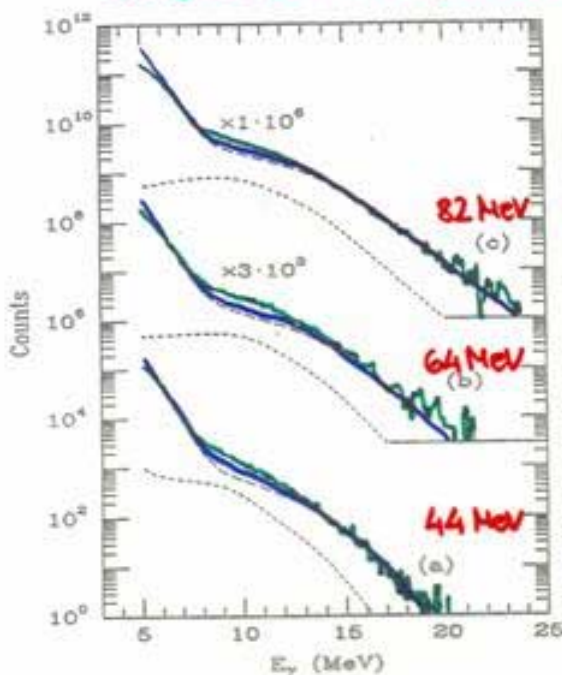
Diószegi J. et al., Phys. Rev. C 46 (1992) 627.

- Experimental results¹ -

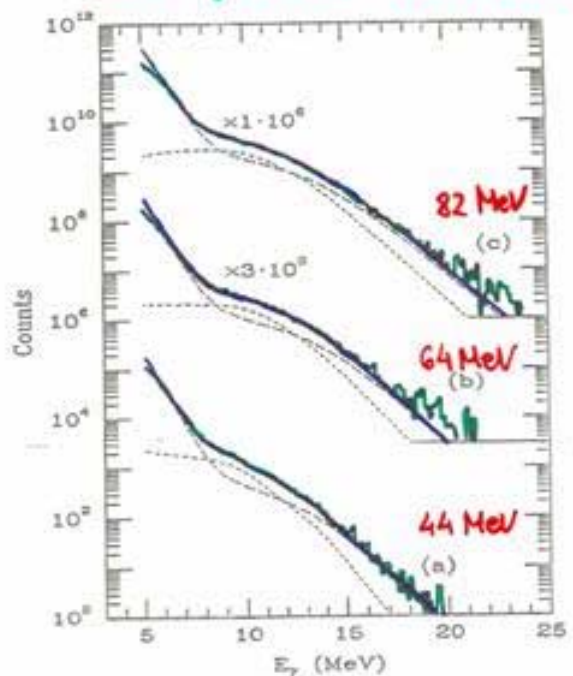
* Thoennessen M. et al., Phys. Rev. Lett. 59 (1987) 2861.

$^{18}\text{O} + ^{208}\text{Pb} \rightarrow \text{Th}$ at $E^* = 44, 64, 82 \text{ MeV}$

no fission hindrance



with fission hindrance

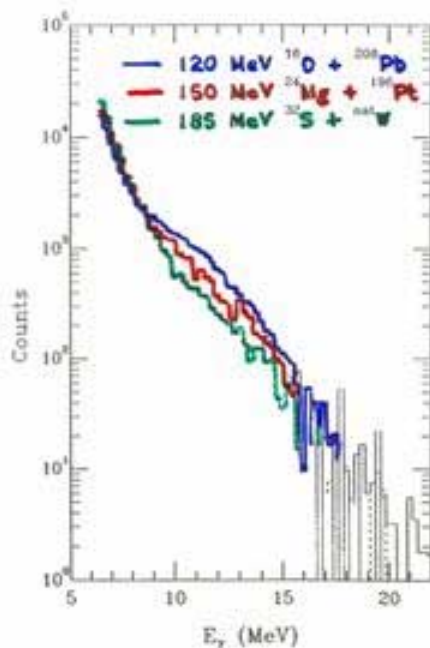
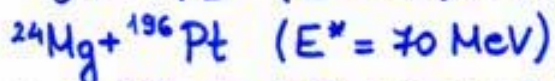
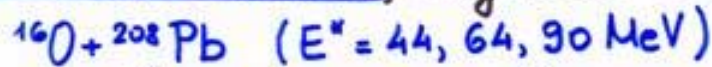


* the best agreement with experimental data :

$$\beta = 5 \cdot 10^{24} \text{ s}^{-1}$$

- Experimental results -

* Butsch R. et al., Phys. Rev. C44 (1991) 1515

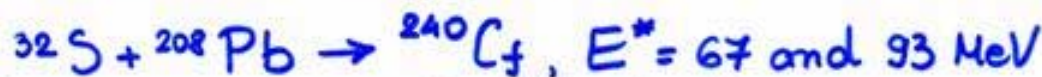


* fusion-fission: $\beta = (10-20) \cdot 10^{24} \text{ s}^{-1} \rightarrow \tau_{\text{Fiss}}^{\text{TOTAL}} \approx 2.9 \cdot 10^{-19} \text{ s}$
 GDR γ rays are mostly emitted before the system passes over the saddle point.

* quasifission: $\beta = 20 \cdot 10^{24} \text{ s}^{-1} \rightarrow \tau_{\text{qf}} = (2-9) \cdot 10^{-20} \text{ s}$;
 $\tau_{\text{qf}} = (2-4) \cdot 10^{-20} \text{ s}$ in Diószegi J. et al., Phys. Rev. C46 (1992) 627 for $^{32}\text{S} + ^{\text{nat}}\text{W}$ at $E^* = 72, 97, 110 \text{ MeV}$

- Experimental results -

* Hofman D.J. et al, Phys. Rev. Lett. 72 (1994) 470.



* at $E^* = 67 \text{ MeV}$ no fission hindrance

* at $E^* = 93 \text{ MeV}$ $\tau_{\text{ssc}} = 30 \cdot 10^{-24} \text{ s}$ ($\beta_0 = 10 \cdot 10^{24} \text{ s}^{-1}$)

with taking $\beta_i = (10-20) \cdot 10^{24} \text{ s}^{-1}$

* Shaw N.P. et al, Phys. Rev. C 61 (2000) 044612

same system extended to higher excitation energies ($E_{\text{max}}^* = 138 \text{ MeV}$)

* best agreement with data by taking $\beta_i = 4 \cdot 10^{24} \text{ s}^{-1}$ and $\beta_0 = 20 \cdot 10^{24} \text{ s}^{-1}$

* the extracted nuclear dissipation coefficient is found to be independent of temperature.

- Temperature / deformation dependence of dissipation -

* Hofman D.J. et al, Phys. Rev. C 51 (1995) 2597.

$^{16}\text{O} + ^{208}\text{Pb}$ at $E^* = 47, 65$ and 84 MeV
simultaneous fit of γ ray spectra, neutrons multiplicities and evaporation residue cross sections $\Rightarrow$
 $\gamma = 16.49 \cdot T_{\text{saddle}} - 18.57$
 $\gamma = 5.71 \cdot T_{\text{saddle}}^2 - 6.81$

* Diószegi J. et al, Phys. Rev. C 61 (2000) 024613
same system at two more E^* (102, 118 MeV)

1. new level density approach (Ignatyuk - Reisdorf)

2. $\gamma(T)$ during the deexcitation process

$$\Rightarrow \gamma = 0.2 + 5T \text{ and } \gamma = 0.2 + 3T^2$$

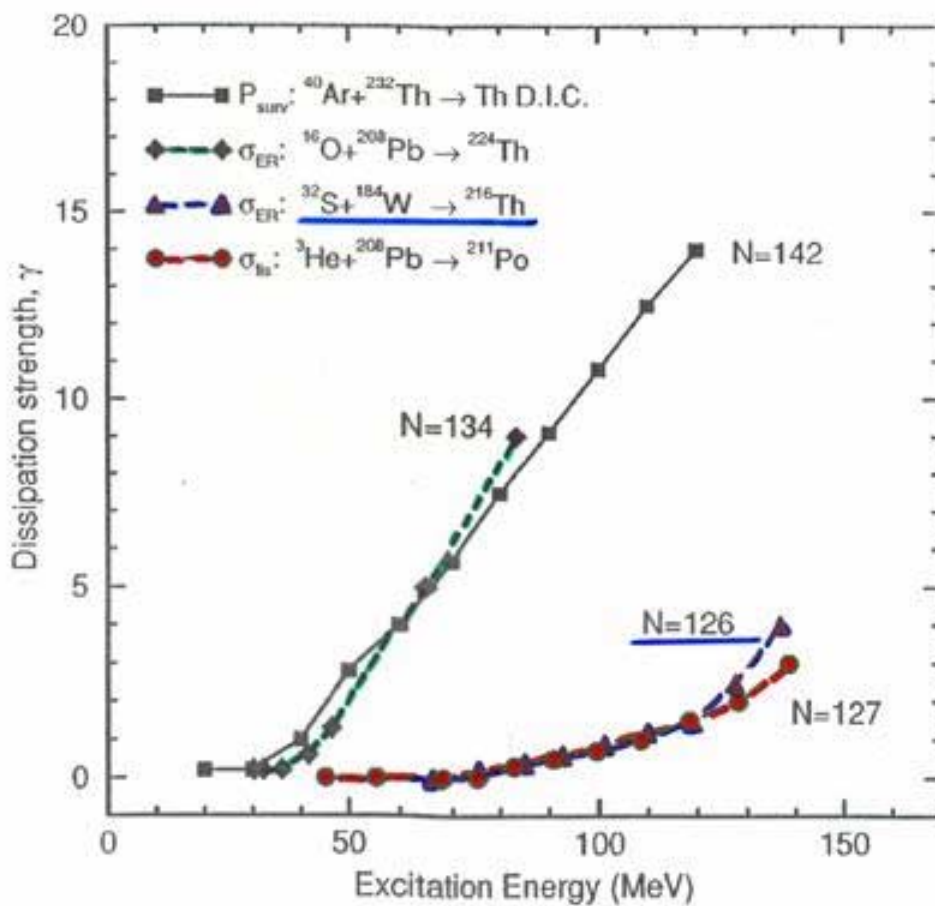
3. additional temperature dependent factor in level density

$$\Rightarrow \gamma = 0.2 + 1.7T^2$$

4. data also fitted with constant $\gamma_i = 2$ inside the saddle and constant $\gamma_o = 10$ outside the saddle.

No definite conclusion!

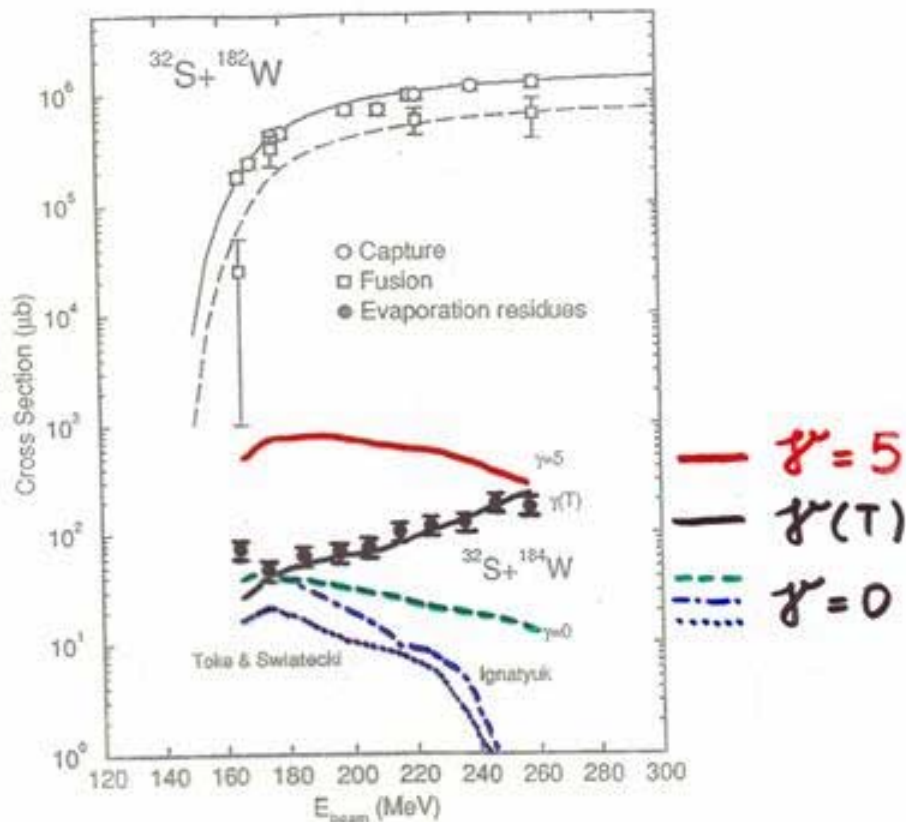
Back B.B. et al, Phys. Rev. C 60 (1999) 044602



- Evaporation residue measurements -

* Back B.B. et al, Phys. Rev. C 60 (1999) 044602

$^{32}\text{S} + ^{184}\text{W}$ at $E_{\text{beam}} = 165 - 257 \text{ MeV}$



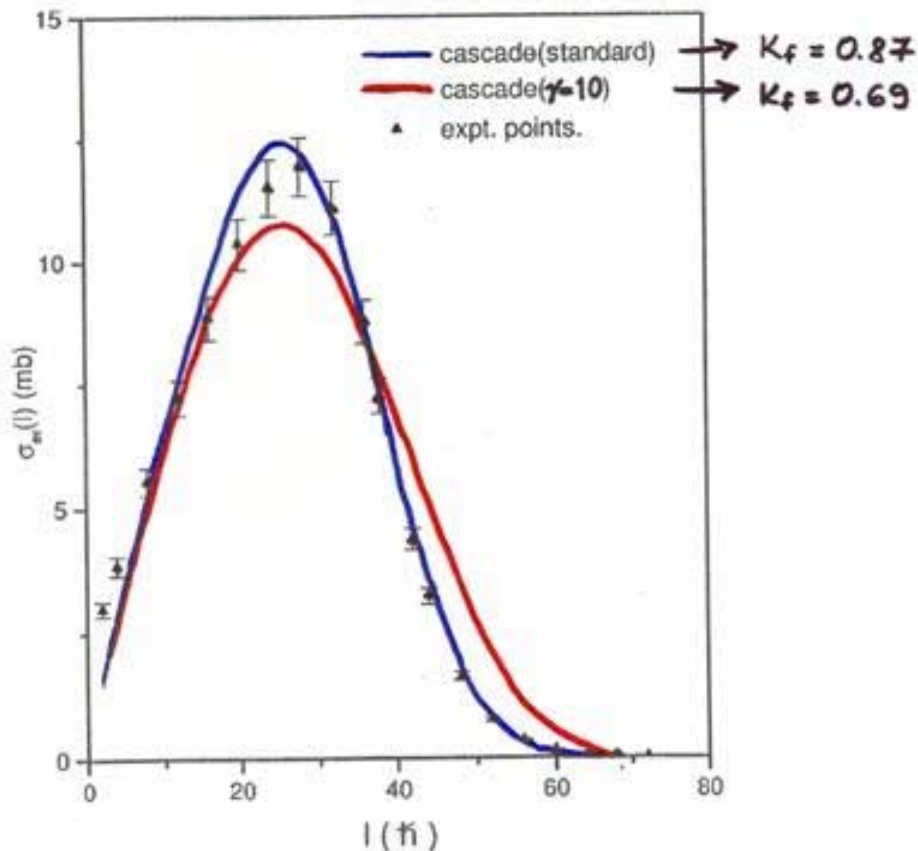
* the dissipation strength is NOT allowed to vary as the system cools down during the particle evaporation cascade.

- Evaporation residue measurements -

* Hui Y.H. et al, Phys. Rev. C 62 (2000) 054604

$^{19}\text{F} + ^{175}\text{Lu}$ at $E_{\text{beam}} = 90 - 125 \text{ MeV}$

* measured ER cross section is not sensitive enough, but the spin dependence is.

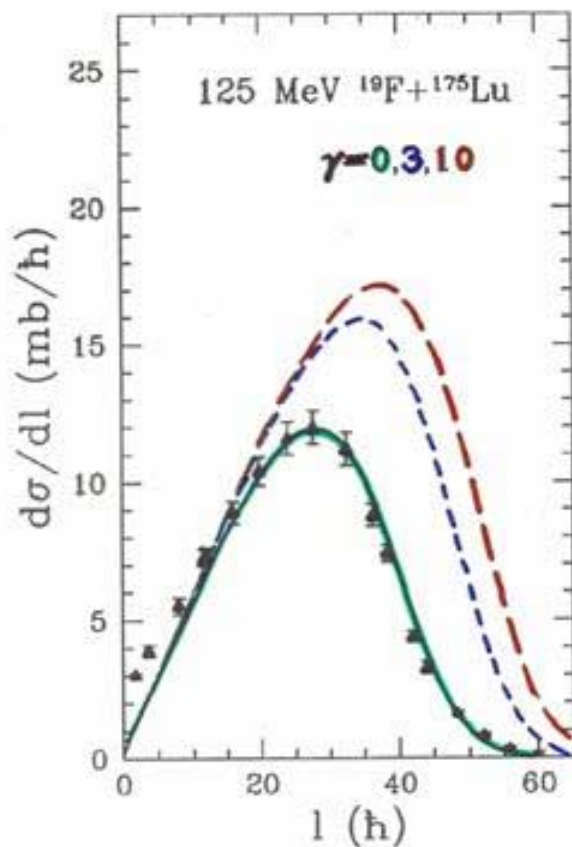


BUT! $\Rightarrow$

* Diószegi J., Phys. Rev. C 64(2001) 019801

* ER cross section is a very sensitive measure of the nuclear viscosity, if the calculations are done without scaling the fission barrier using the appropriate level density method:

Ignatyuk-Reisdorf approach $\Rightarrow$



- Level density calculations -

* original version of CASCADE:

$E^* < 10 \text{ MeV} \rightarrow$ parametrization of
Dilg W. et al, Nucl. Phys.
A214 (1973) 269.

$E^* > 20 \text{ MeV} \rightarrow$ nucleus $\leftrightarrow$ liquid drop

* improved excitation energy dependent description:

- Ignatyuk A.V. et al, Sov. J. Phys. 21
(1975) 255

$$a(u) = \tilde{a}(A) \left[1 + \frac{\delta W}{u} (1 - e^{-u/E_d}) \right]$$

- Reisdorf W., Z. Phys. A300 (1981) 227.

$$\tilde{a}(A) = 0.04543 r_0^3 + 0.1355 r_0^2 A^{-1/3} B_0 + 0.1426 r_0 A^{-2/3} B_k$$

* additional temperature dependence:

$$a(T) = a(u) \cdot [1 - k \cdot f(T)]$$

$$f(T) = 1 - \exp \left[- (T A^{1/3} / 21)^2 \right]$$

$$k = 0.4$$

(Ghulom Y. and Natorwitz J.B., Phys. Rev. 44
(1991) 2878)

- 1-exp versus step -

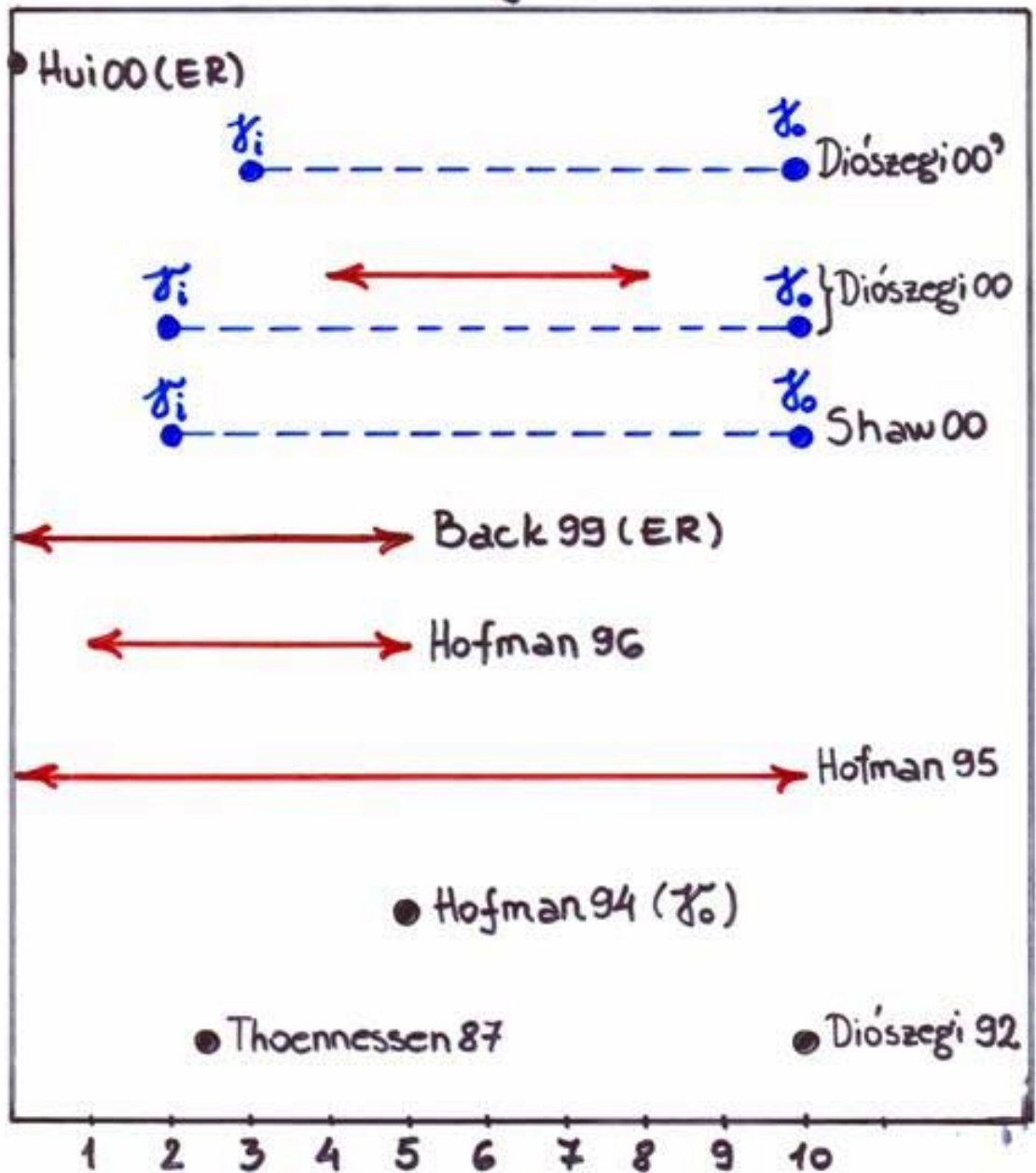
Γ_{KRAMERS}
 $\downarrow$

$\Gamma_{\text{f}}^{\text{BW}}$
 $\downarrow$

System	1-exp	step
$^{32}\text{S} + \text{W}$	$\bar{\tau}_{\text{f}} \approx 10 \cdot 10^{-24} \text{ s}^{(4)}$ $\bar{\tau}_{\text{QF}} = (20-40) 10^{-24} \text{ s}^{(2)}$	$\bar{\tau}_{\text{PRE}} \sim 10 \cdot 10^{-24} \text{ s}^{(3)}$ $\bar{\tau}_{\text{POST}} \sim 20 \cdot 10^{-24} \text{ s}$
$^{32}\text{S} + ^{208}\text{Pb}$	$\bar{\tau}_{\text{SSE}} = 30 \cdot 10^{-24} \text{ s}^{(4)}$ $\bar{\tau}_{\text{SSE}} = 54 \cdot 10^{-24} \text{ s}^{(5)}$	$\bar{\tau}_{\text{PRE}} \sim 10 \cdot 10^{-24} \text{ s}^{(3)}$ $\bar{\tau}_{\text{POST}} \sim 20 \cdot 10^{-24} \text{ s}$

- (1) Back B.B. et al, Phys. Rev. C (1999) 044602
 $^{32}\text{S} + ^{184}\text{W}$, ER
- (2) Dio'szegi J. et al, Phys. Rev. C 46 (1992) 627
 $^{32}\text{S} + ^{\text{nat}}\text{W}$, GDR clock
- (3) Morton C.R. et al, J. Phys. G 23 (1997) 1383.
 $^{32}\text{S} + ^{\text{nat}}\text{W}$, GDR clock
- (4) Hofman D.J. et al, Phys. Rev. Lett. 72 (1994) 470.
- (5) Phaw N.P. et al, Phys. Rev. C 61 (2000) 024613.

- Summary -



$$\gamma = \frac{\beta}{2\omega} = \frac{\eta}{2H\omega}$$

— temperature dependence
 --- deformation dependence

1. Thoennessen M. et al, Phys. Rev. Lett. 59(1987)
 $^{16}\text{O} + ^{208}\text{Pb}$, $E^* = 44, 64, 82 \text{ MeV}$
2. Diószegi J. et al, Phys. Rev. C 46(1992)627
 $^{16}\text{O} + ^{208}\text{Pb}$, $E^* = 82 \text{ MeV}$
 $^{32}\text{S} + ^{nat}\text{W}$, $E^* = 72 - 110 \text{ MeV}$
3. Hofman D.J. et al, Phys. Rev. Lett. 72(1994)470
 $^{32}\text{S} + ^{208}\text{Pb}$ $E^* = 67 - 93 \text{ MeV}$
4. Hofman D.J. et al, Phys. Rev. C 51(1995)2597
 $^{16}\text{O} + ^{208}\text{Pb}$ $E_{\text{lab}} = 80 - 140 \text{ MeV}$
5. Hofman D.J. et al, Nucl. Phys. A 599(1996)23c
 $^{32}\text{S} + ^{nat}\text{W}$ $E^* = 50 - 140 \text{ MeV}$
6. Back B.B. et al, Phys. Rev. C 60(1999)044602
 $^{32}\text{S} + ^{184}\text{W} \rightarrow \text{ER}$, $E_{\text{lab}} = 165 - 257 \text{ MeV}$
7. Shaw J.P. et al, Phys. Rev. C 61(2000)044612
 $^{32}\text{S} + ^{208}\text{Pb}$ $E^* = 46 - 138 \text{ MeV}$
8. Diószegi J. et al, Phys. Rev. C 61(2000)024613
 $^{16}\text{O} + ^{208}\text{Pb}$ $E^* = 46 - 118 \text{ MeV}$
9. Diószegi J. et al, Phys. Rev. C 63(2000)014611
 $^{19}\text{F} + ^{231}\text{Ta}$ $E^* = 121, 139 \text{ MeV}$
10. Hui Y.K. et al, Phys. Rev. C 62(2000)054604
 $^{19}\text{F} + ^{175}\text{Lu} \rightarrow \text{ER}$ $E^* = 90 - 125 \text{ MeV}$